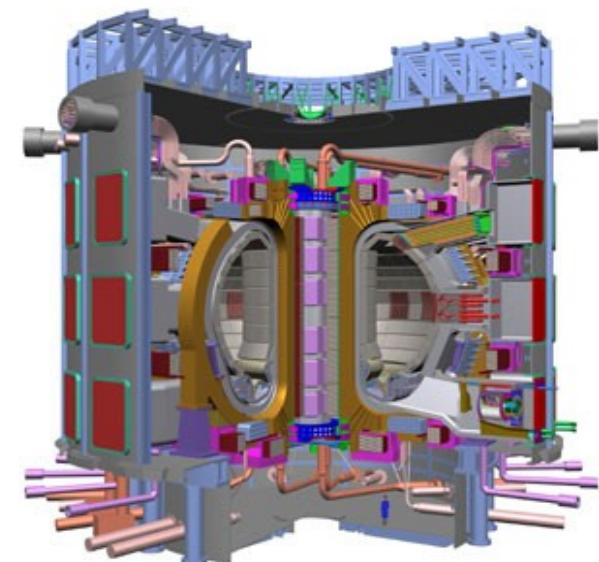


Atomic structure and dynamics

-- need and requirements for accurate atomic calculations

- Analysis and interpretation of optical and x-ray spectra (astro physics)
- Isotope shifts and hyperfine structures
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- Parity nonconservation (PNC)
- Search for electric dipole moments



Relativistic effects in heavy and superheavy elements

... to be considered in accurate atomic calculations

S. Fritzsche, MPI-K and GSI Darmstadt

11th October 2007

Atomic interactions are well known:

- QED as the well established basis
- Atomic shell modell
 - $\psi(r,\theta,\phi) = R_{nl}(r) Y_{lm}(\theta,\phi)$,
 - „aufbau principle“: Successive filling of subshells and shells

Outline of this talk:

- i) „Electronic correlations“: The challenge of open shells
- ii) (Super-) Heavy elements: Rapid increase of relativity
- iii) Multiphoton processes in high-Z systems
- iv) Nonradiative transitions and autoionization

„Electronic correlations“

-- Fine-structure of open-shell configurations

Dimension of the
Hilbert space

p^4	5
$p^3 s$	10

$p^3 s$

p^4

„Electronic correlations“

-- Fine-structure of open-shell configurations

Dimension of the Hilbert space

p^4 5
 $p^3 s$ 10

d^8 9
 $d^7 p$ 110

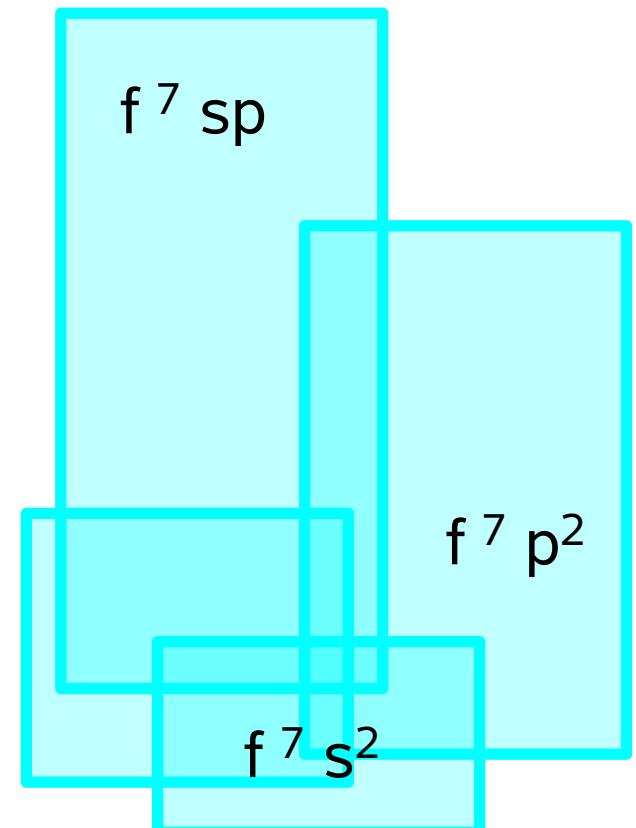
$f^7 s^2$ 327
 $f^7 sp$ 3808
 $f^7 p^2$ 4724
 $f^6 spd$ 31804

$p^3 s$

p^4

$\sim 1 \dots 3 \%$
level and transition energies
 $\sim 10 \dots 30 \%$

Computational requirements depend very critically on the shell structure of the atoms and ions !



„Electronic correlations“

-- Fine-structure of open-shell configurations

Dimension of the Hilbert space

p^4
 $p^3 s$

5
10

$p^3 s$

d^8
 $d^7 p$

9
110

p^4

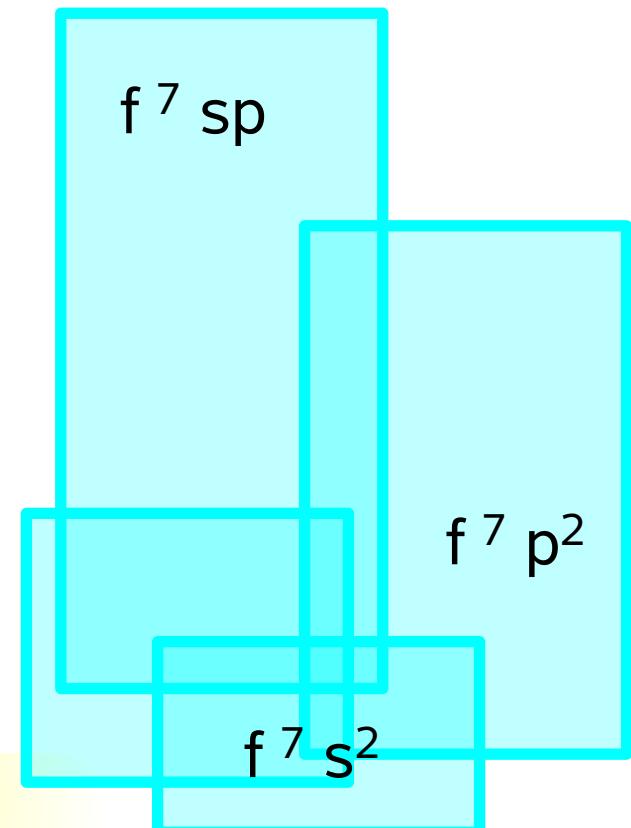
Computational requirements depend very critically on the shell structure of the atoms and ions !

$f^7 s^2$	327
$f^7 sp$	3808
$f^7 p^2$	4724
$f^6 spd$	31804

~ 1 .. 3 %

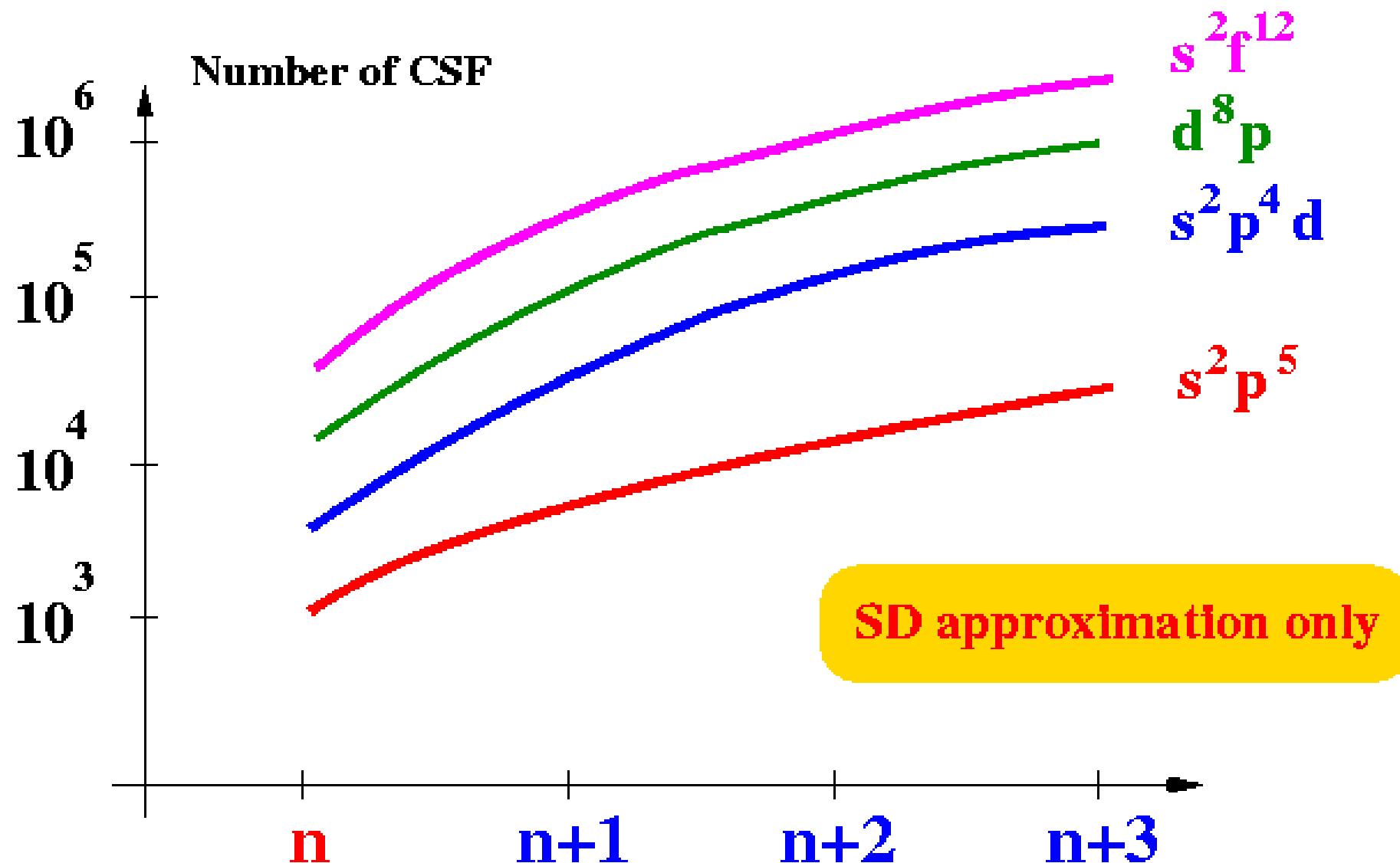
level and transition energies

~ 10 ... 30 %



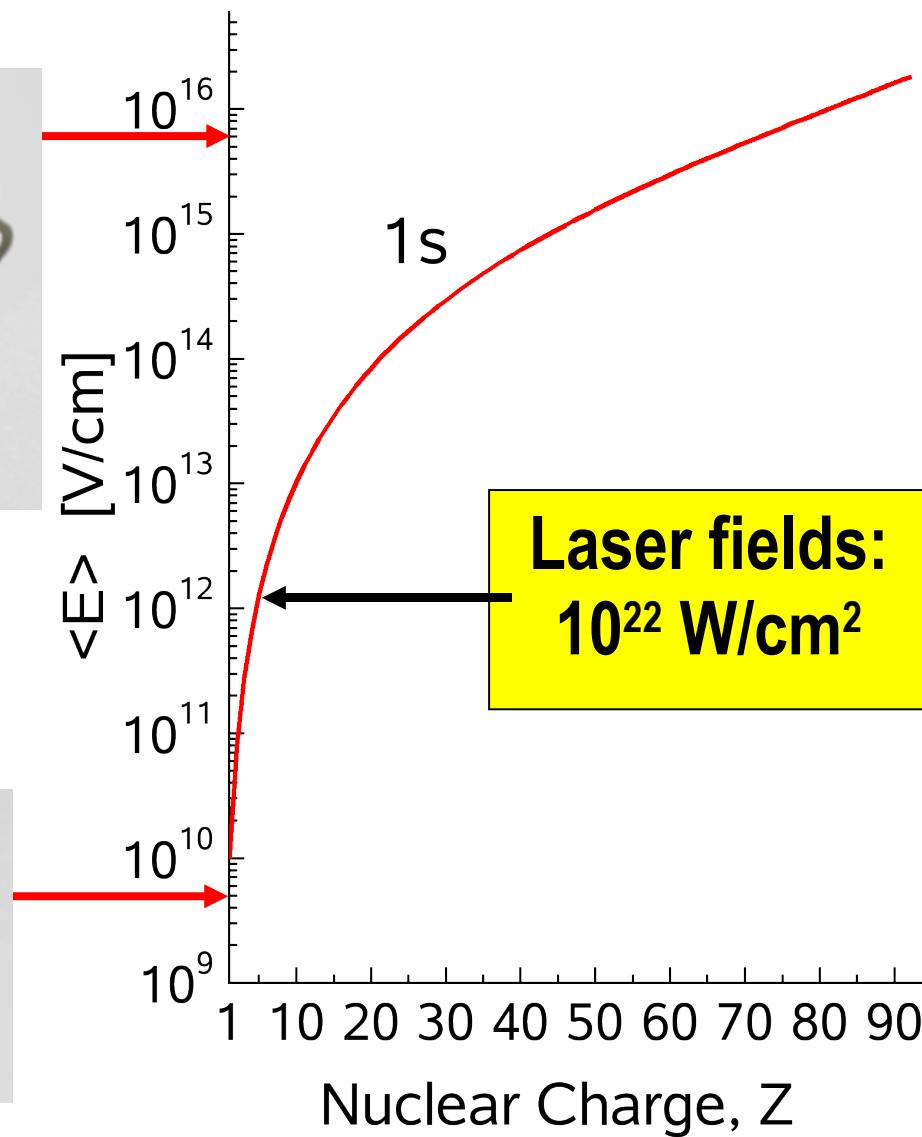
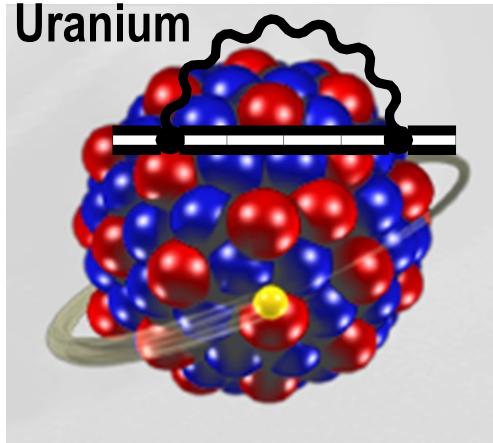
Concept of electron configuration gets lost !

Wave function (CSF) expansions for open-shell structures



Extreme Static Electromagnetic Fields

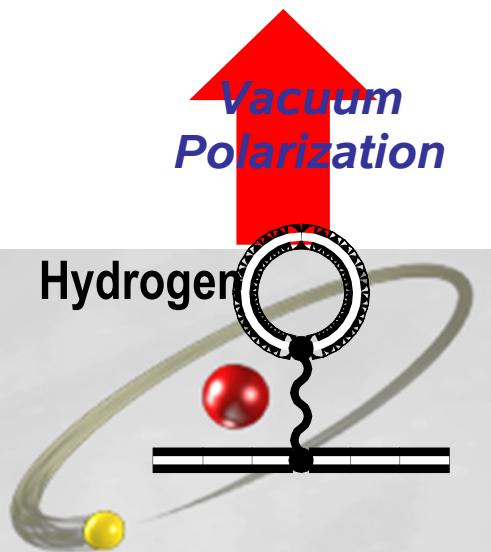
Self Energy



$$\Delta E \approx 500 \text{ eV}$$
$$Z \cdot \alpha \approx 1$$



Quantum
Electro-
Dynamics

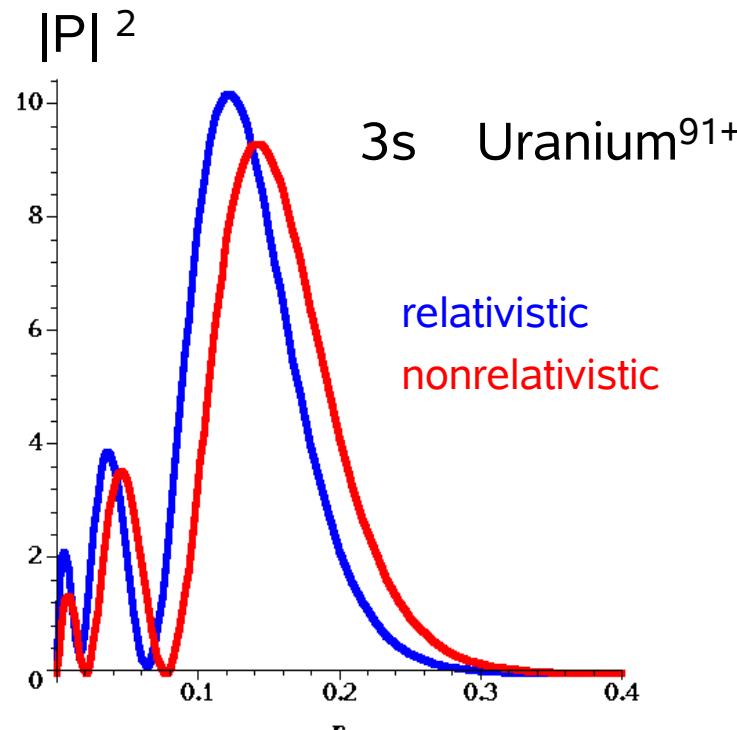


$$\Delta E \approx 10^{-6} \text{ eV}$$
$$Z \cdot \alpha \approx 10^{-2}$$



Relativistic and quantum-electrodynamical corrections

-- Test of QED in hydrogen-like uranium



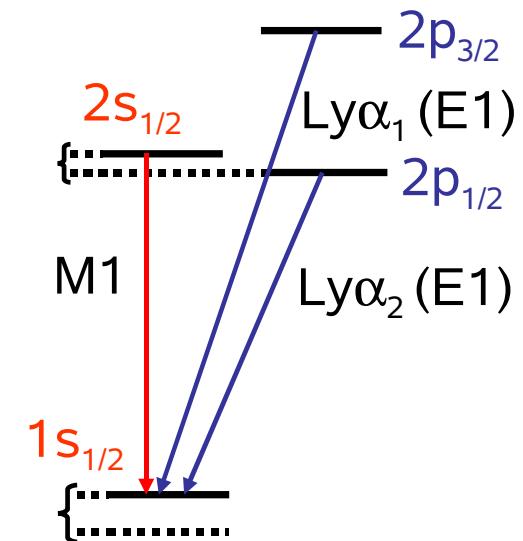
Relativistic contraction
of the wave functions

Use of the one-particle Dirac operator

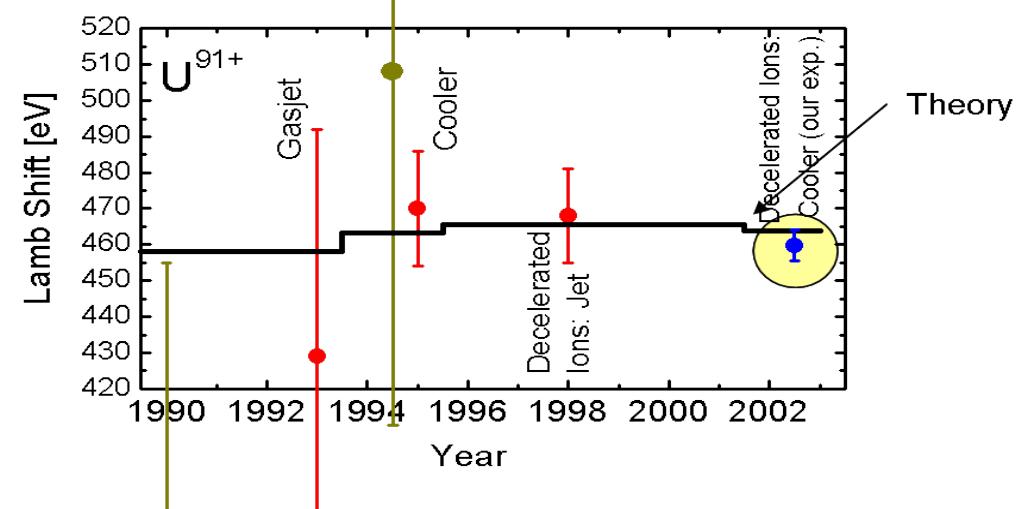
1s-Lamb Shift

Experiment: $459.8 \text{ eV} \pm 4.6 \text{ eV}$

Theory: 463.95 eV

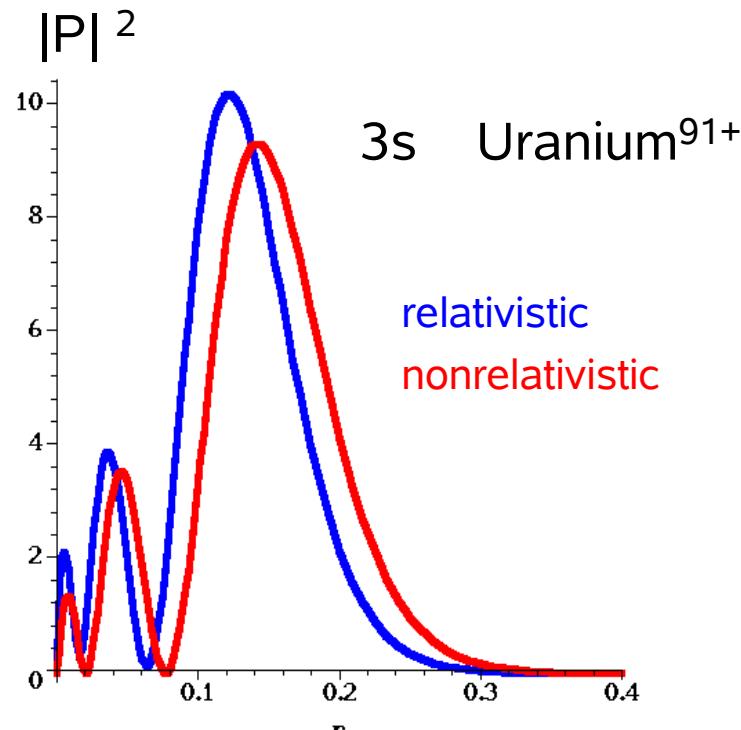


A. Gumberidze, PhD thesis (2003),
PRL 94 (2005) 223001.



Relativistic and quantum-electrodynamical corrections

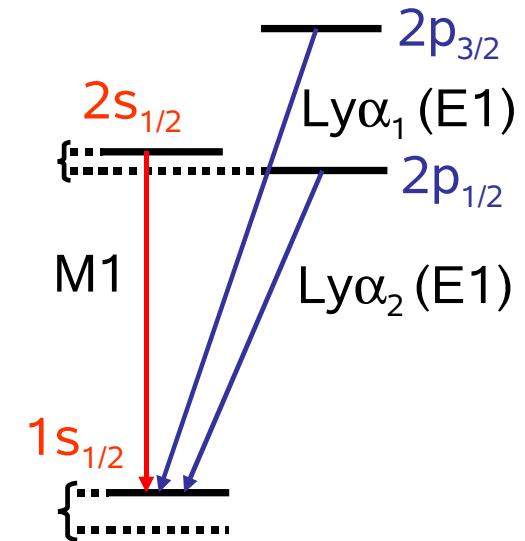
-- Test of QED in hydrogen-like uranium



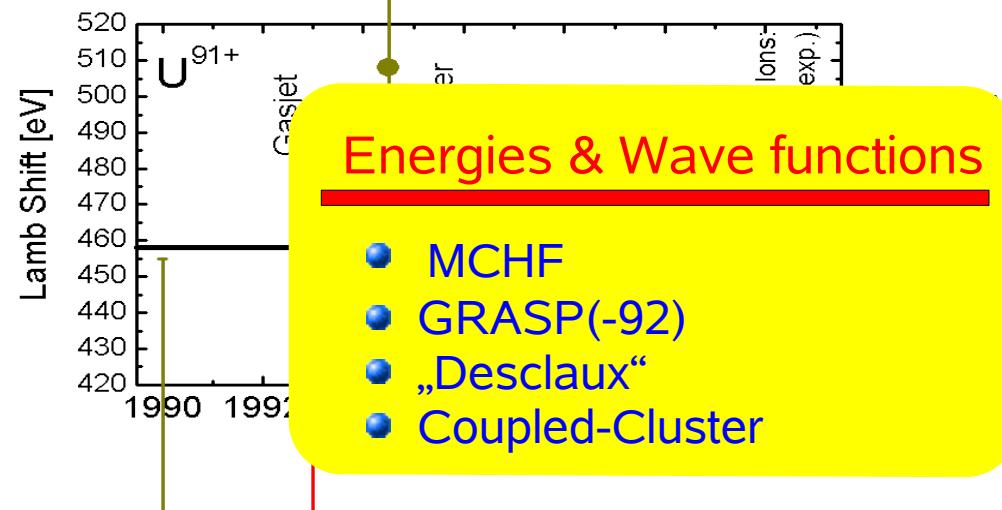
Relativistic contraction
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RATIP

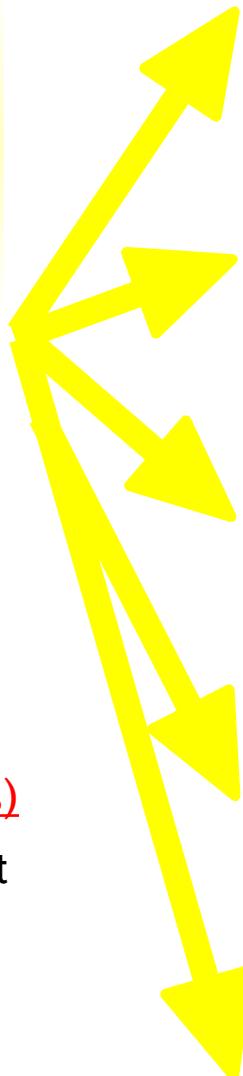
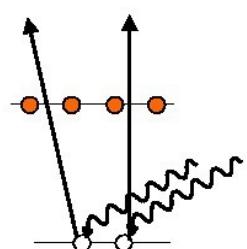
Relativistic Atomic Transition and Ionization Properties

(CPC library)

$$\psi_{\alpha}(PJM) = \sum_r^{n_c} c_r(\alpha) |\gamma_r PJM\rangle$$

Many-electron basis (wave function expansions)

- Construction and classification of N-particle Hilbert spaces
- Shell model: Systematically enlarged CSF basis
- Interactions
 - Dirac-Coulomb Hamiltonian
 - Breit interactions + QED
 - Electron continuum; scattering phases
- Coherence transfer and Rydberg dynamics



Relativistic CI wave functions
including QED estimates and
mass polarization

RELCI, CPC 148 (2002) 103

LSJ spectroscopic notation
from jj-coupled
computations

LSJ, CPC 157 (2003) 239

Auger rates, angular distribu-
tions and spin polarization;
level widths

AUGER

Photoionization cross sect-
ions and (non-dipole) angular
parameters

PHOTO

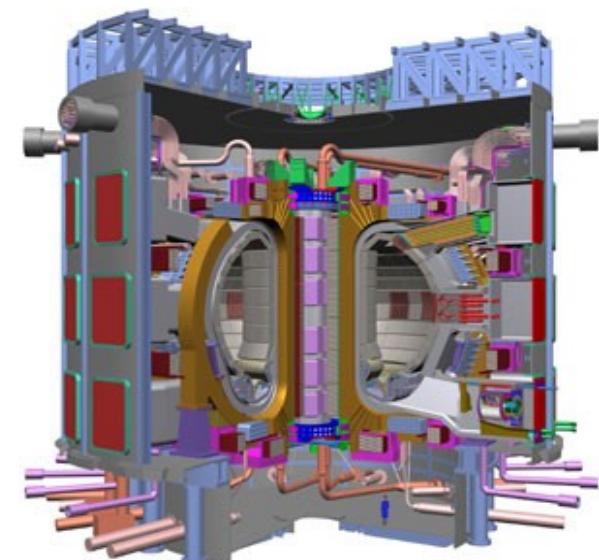
Radiative and dielectronic
recombination; angle-angle
correlations

...

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Systematic multiconfiguration calculations

(CI, MCHF, MCDF)

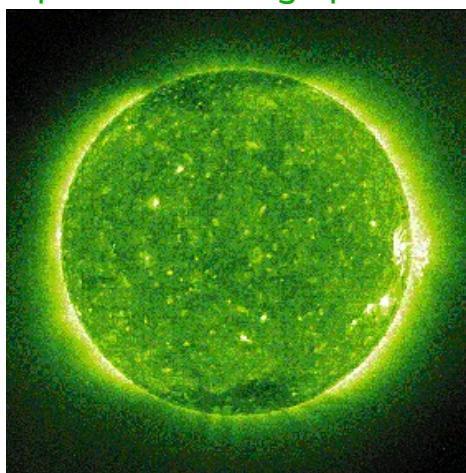
Up to now:

- Term- and hyperfine structure for light elements ($Z \leq 28$)
- Resonance and intercombination lines
- Lifetimes

Benchmarks:
He, Li-like, C²⁺

Example : EUV spectra of multiple-charged iron from the sun

Spectra involving open d-shells



Iron is one of the most abundant heavy elements in the universe (opacity project)

Fe X ... XIV 3s 3p n+1, 3s² 3p n-1 3d

$$\Delta E / E < 1\%$$

$$\Delta A / A = 5 .. 20 \%$$

Improved by factor 5 !

- Line identification
- Improved level structure
- Reliable lifetimes

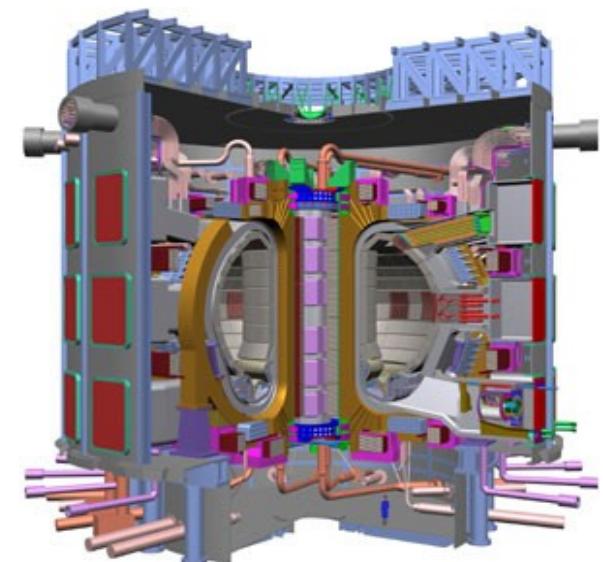
Fe X: 31 low-lying levels (Dong et al., MNRAS, 1999)

Fe XI: 47 levels (Fritzsche et al., MNRAS, 2000)

Atomic structure and dynamics

-- need and requirements for accurate atomic calculations

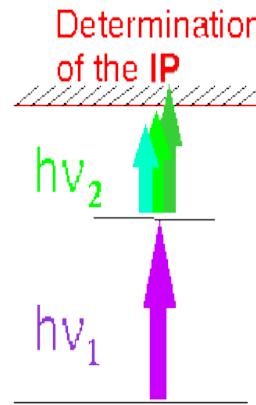
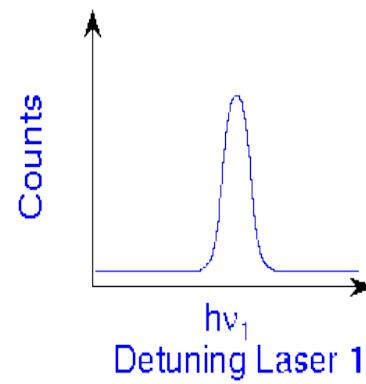
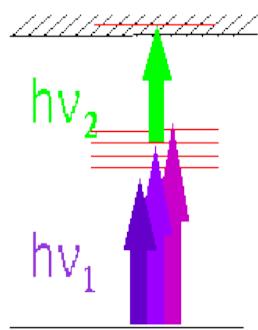
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Optical spectroscopy at Fermium ($Z = 100$)

-- first observation and classification of atomic levels

Determination of hfs
and isotope shifts



Atomic Physics:

- Atomic Structure

- Ionization potentials

Nuclear Physics:

- Nuclear spins

- Moments

- Changes of charge radii

$5f^{12} 7s^2$, $JP = 6^+$

$5f^{12} 7s 7p$, $JP = 6^-, 5^-$

$5f^{12} 7s 7p$, $JP = 6^-, 5^-, 7^-$ $^3H_{6,5,7}^o$ (?)

$^3H_6^e$

$^5G_{6,5}^o$

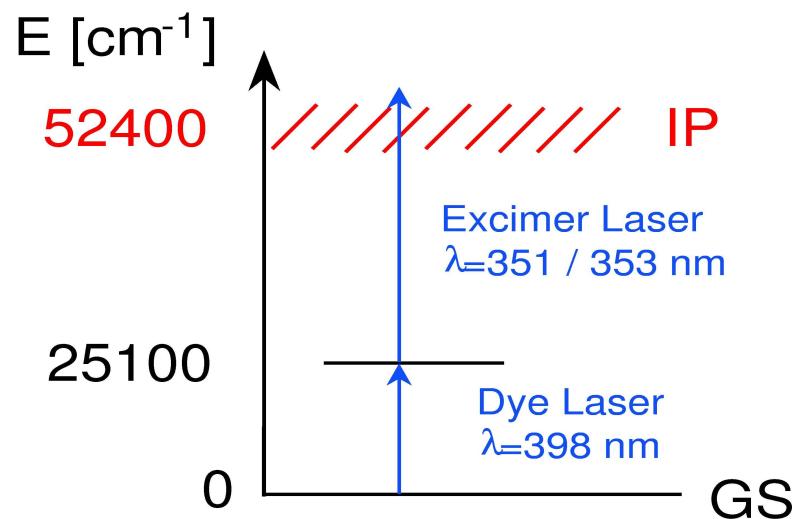


TABLE I. Results of MCDF calculations. Accuracy of transition energy $\bar{\nu}$ is $\Delta \bar{\nu} = \pm 2400 \text{ cm}^{-1}$, A_{ki} = Einstein coefficient, classification according to the largest coefficient c in the CSF expansion.

No.	$\bar{\nu}$ (cm^{-1})	J	A_{ki} s^{-1}	Config.	Term	$ c ^2$
1	0	6	0	$5f^{12}7s^2$	$^3H_6^e$	0.96
2	25 226	6	1.89×10^6	$5f^{12}7s7p$	$^5I_6^o$	0.46
3	25 471	5	1.28×10^6	$5f^{12}7s7p$	$^5G_5^o$	0.34
4	27 394	6	2.43×10^8	$5f^{12}7s7p$	$^3H_6^o$	0.62
5	27 633	5	1.98×10^8	$5f^{12}7s7p$	$^3G_5^o$	0.60
6	27 802	7	3.67×10^8	$5f^{12}7s7p$	$^3I_7^o$	0.66

Low-lying resonances for heavy and super-heavy elements

... for lutetium ($Z=71$) and lawrencium ($Z=103$)

TABLE I. The transition energies in cm^{-1} of $nd \ ^2D_{3/2} - (n + 1)p \ ^2P_{1/2,3/2}^o$ and the size of CSF expansions for Lu ($n = 5$) and Lr ($n = 6$).

Expansion	$^2D_{3/2} - ^2P_{1/2}^o$	$^2D_{3/2} - ^2P_{3/2}^o$	CSF ($^2D_{3/2}/^2P_{1/2}^o/^2P_{3/2}^o$)
Lu			
VV + CV($4f^{14}$)	3989	7276	4354/2071/3813
VV + CV($5p^64f^{14}$)	8004	11483	5600/2764/5073
VV + [(CV + CC)($5p^64f^{14}$)]	3857	7130	128 763/36 974/100 277
VV + [(CV + CC)($4d^{10}5s^25p^64f^{14}$)]	4186	7462	305 717/87 241/236 554
RCC [7]	3828	7140	
DFT [10]	3862		
Exp.	4136	7476	
DHF Breit Correction	87	53	
DHF Breit & QED Correction	76	43	
Lr			
VV + CV($5f^{14}$)	-1298	9137	3659/1842/3338
VV + CV($6p^65f^{14}$)	1339	12 761	4708/2495/4495
VV + [(CV + CC)($6p^65f^{14}$)]	-1953	6469	125 325/37 333/97 500
VV + [(CV + CC)($5d^{10}6s^26p^65f^{14}$)]	-1127	7807	330 252/95 969/246 376
RCC	-1388	6960	
RCC with Breit	-1263	7010	
DHF Breit Correction	97	4	
DHF Breit & QED Correction	59	-26	

Low-lying resonances for heavy and super-heavy elements

-- oscillator strengths in different gauges

TABLE II. The oscillator strengths of $nd \ ^2D_{3/2} - (n + 1)p \ ^2P_{1/2,3/2}^o$ for Lu ($n = 5$) and Lr ($n = 6$).

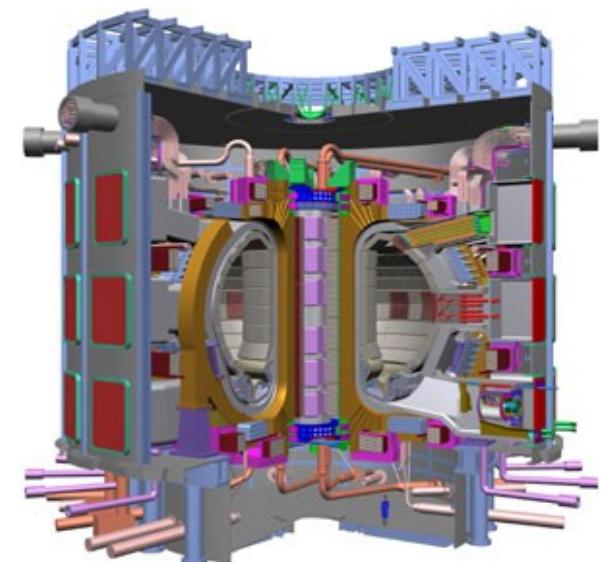
Expansion	$^2D_{3/2} - ^2P_{1/2}^o$			$^2D_{3/2} - ^2P_{3/2}^o$		
	gf_L	gf_V	Scaled gf_L	gf_L	gf_V	Scaled gf_L
Lu						
VV + CV($4f^{14}$)	0.0304	0.0582	0.0315	0.0111	0.0219	0.0114
VV + CV($5p^6 4f^{14}$)	0.0511	0.1552	0.0264	0.0144	0.0467	0.0094
VV + [(CV + CC)($5p^6 4f^{14}$)]	0.0908	0.3835	0.0974	0.0322	0.0856	0.0337
VV + [(CV + CC)($4d^{10} 5s^2 5p^6 4f^{14}$)]	0.1043	0.3345	0.1031	0.0354	0.0742	0.0355
Lr						
VV + CV($5f^{14}$)	-0.0162	-0.0076		0.0210	0.0313	
VV + CV($6p^6 5f^{14}$)	0.0144	0.2359		0.0227	0.0839	
VV + [(CV + CC)($6p^6 5f^{14}$)]	-0.0624	-0.0002		0.0414	0.0867	
VV + [(CV + CC)($5d^{10} 6s^2 6p^6 5f^{14}$)]	-0.0378	-0.0024		0.0519	0.0685	

Good accuracy of the (atomic) energies is
a necessary, but not a sufficient criterion !

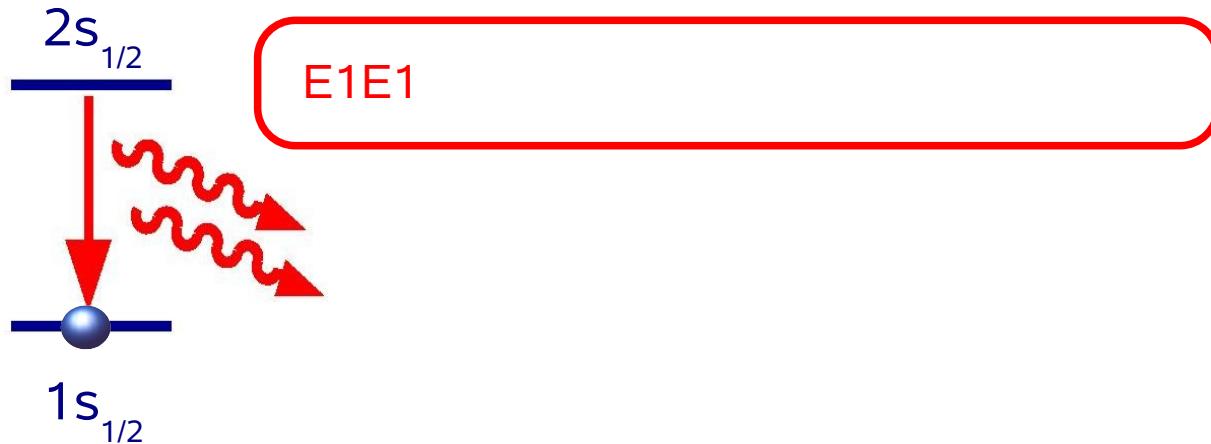
Atomic structure and dynamics

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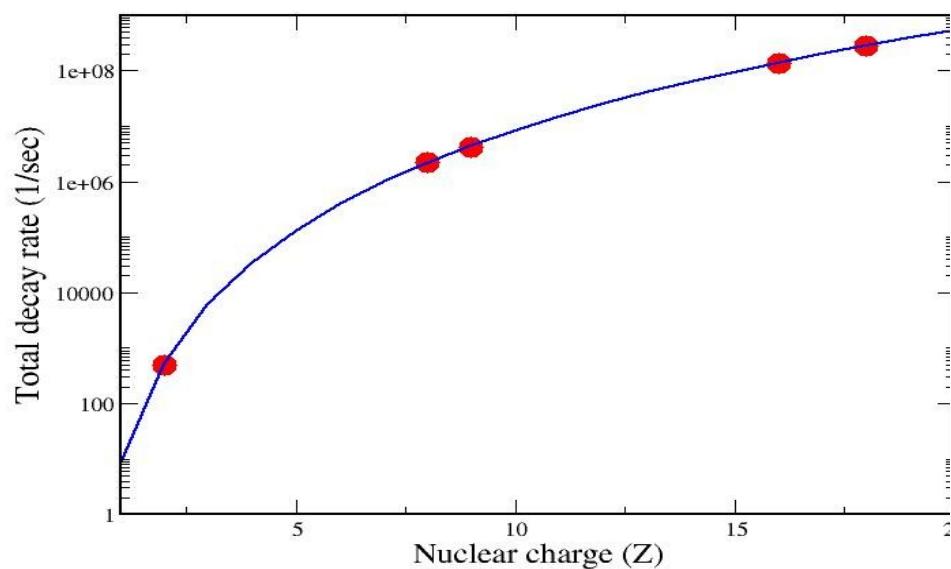
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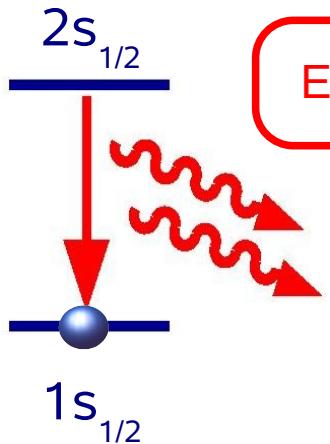
Two-photon decay of highly-charged ions



$$\Gamma_{tot} \approx \Gamma_{E1E1} = 8.229 \cdot Z^6$$



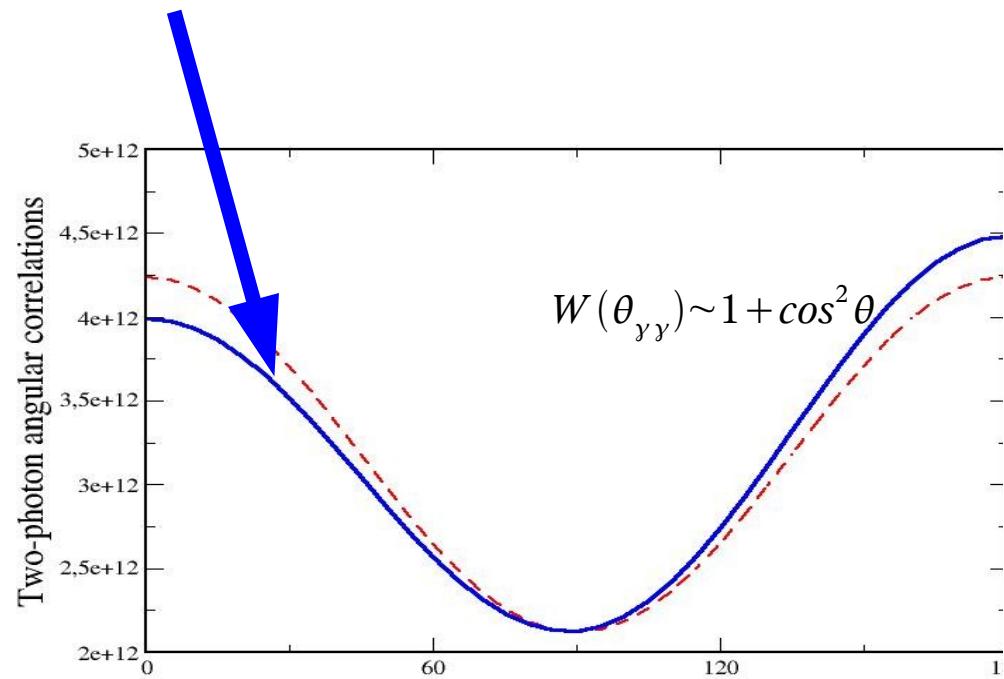
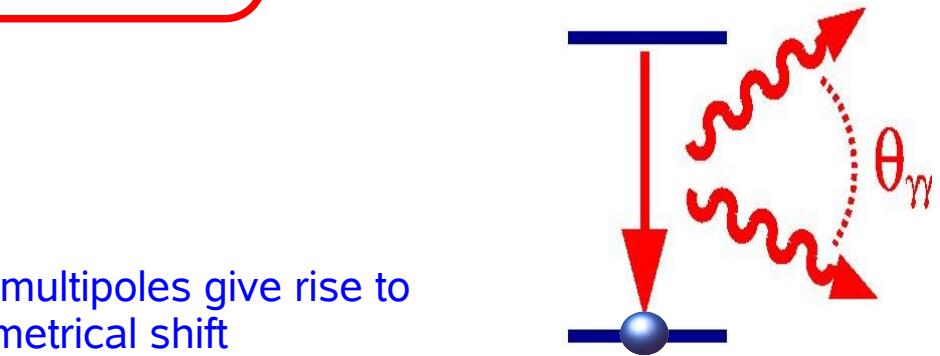
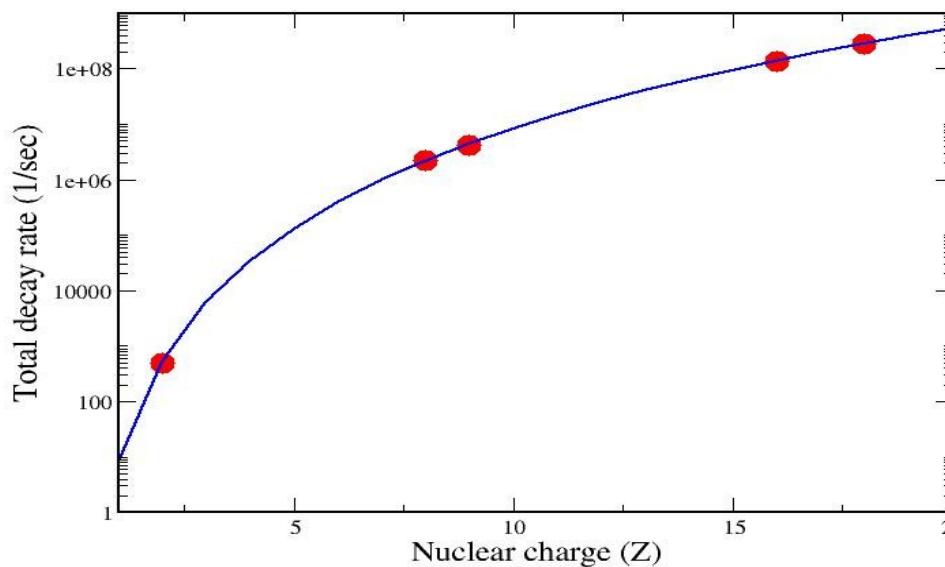
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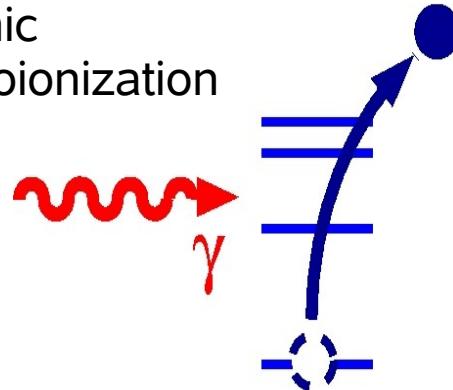
E1E1 + E1M2 + M1M1+E2E2 + E2M1....

Higher multipoles give rise to an asymmetrical shift

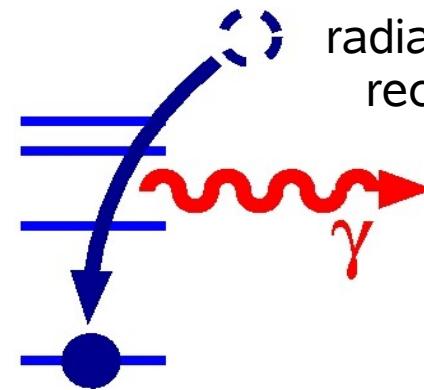
$$\Gamma_{tot} \approx \Gamma_{E1E1} = 8.229 \cdot Z^6$$



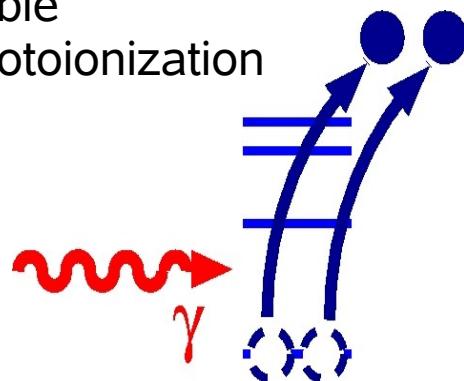
atomic
photoionization



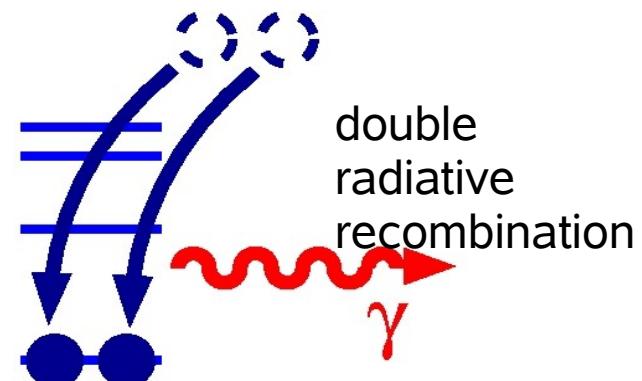
radiative
recombination



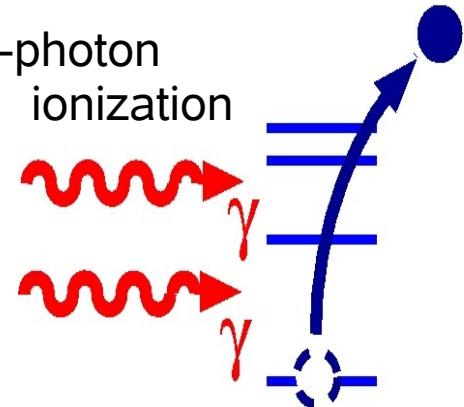
double
photoionization



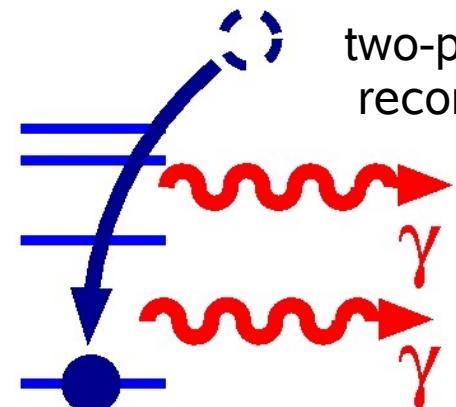
double
radiative
recombination



two-photon
ionization



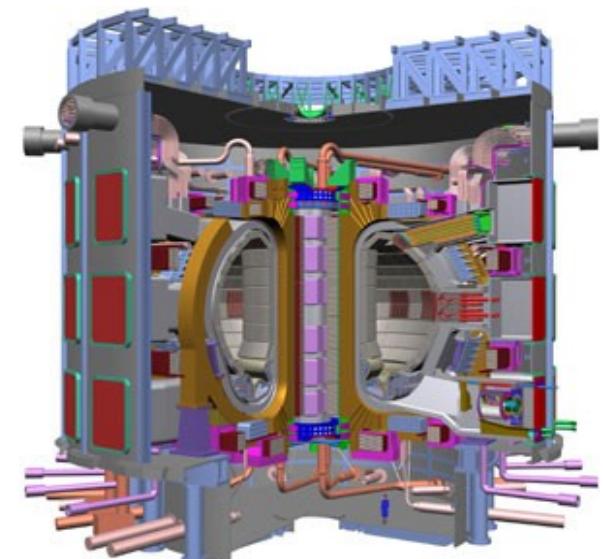
two-photon
recombination



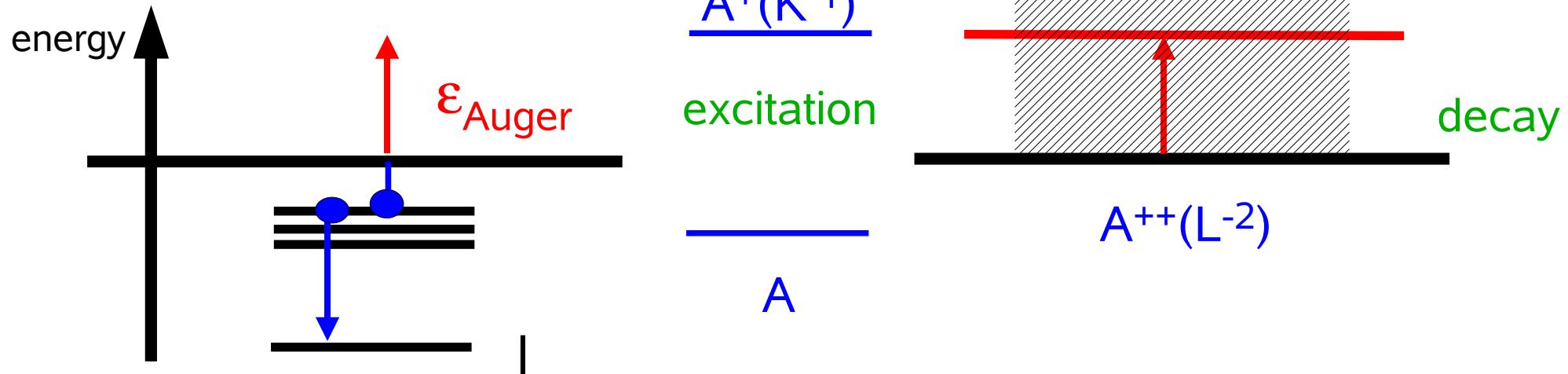
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Auger emission of excited atomic states



$$H = \sum_i (h_i + u(r_i))$$

$$H = \sum_i h_i + \sum_{i < j} \frac{1}{r_{ij}}$$

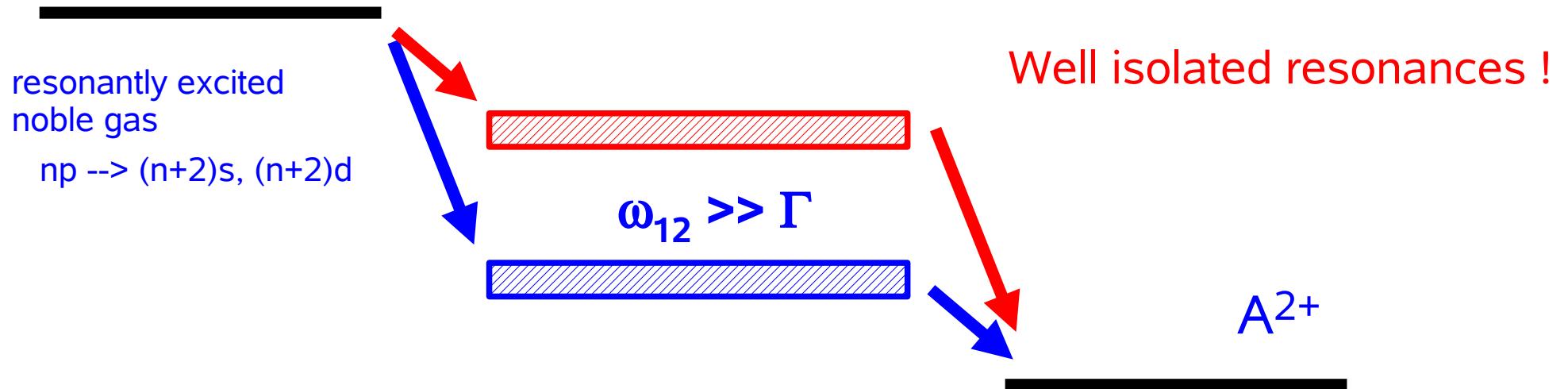
Wentzel's ansatz: Autoionization is caused by electron-electron interactions which cannot be considered in an one-particle picture.

$$\sum_{i < j} \frac{1}{r_{ij}} - \sum_i u(r_i)$$

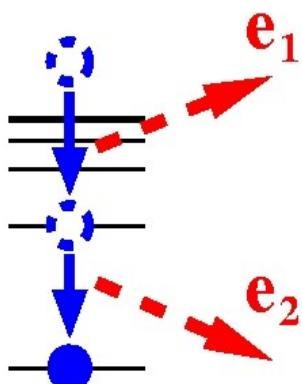
Ideal tool for a better understanding of electronic correlations !

Coherence transfer in the Auger cascades of noble gases

-- a signature of the „atomic double slit“

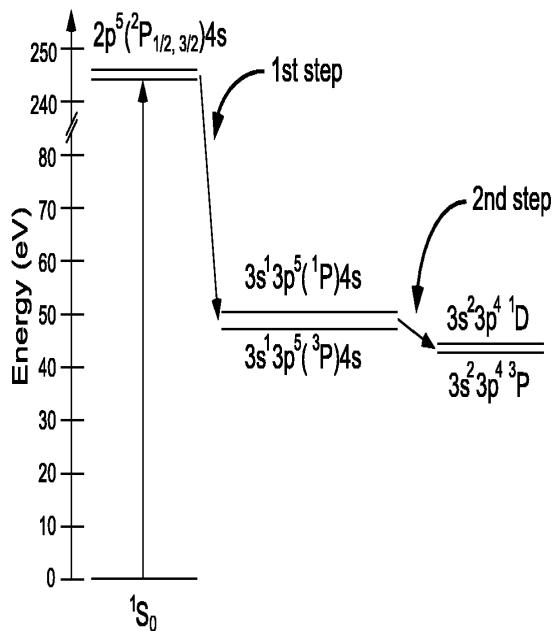


Decay branches are independent; „path“ can be determined by measuring the energy spectrum.

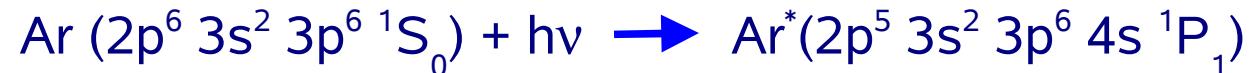


Collaboration with Nicolai Kabachnik (Bielefeld);
experiments by Kyioshi Ueda and coworkers at SPring8, Japan

Excitation and two-step Auger cascades in noble gases



Photoabsorption:



First decay:



Second decay:



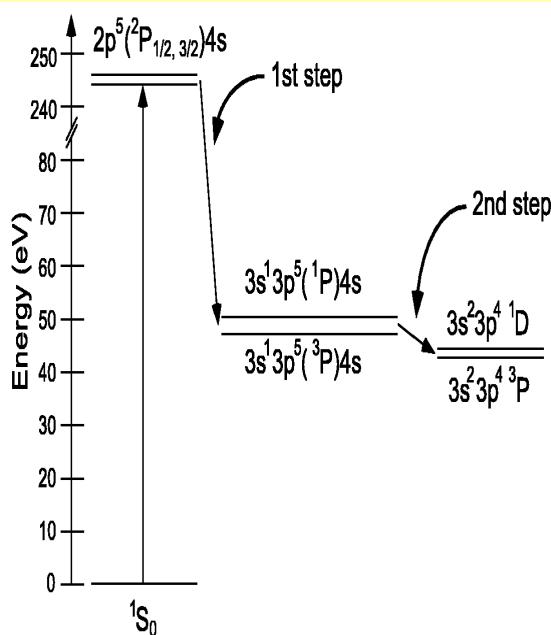
Ne:	500 : 1
Ar:	80 : 1
Kr:	25 : 1
Xe:	8 : 1

$$\frac{A_{\text{resonance}}}{A_{\text{intercombination}}}$$

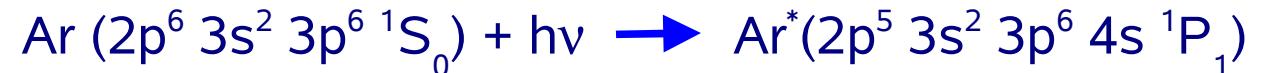


Relativity enters here in two ways !

Excitation and two-step Auger cascades in noble gases



Photoabsorption:



First decay:



Second decay:

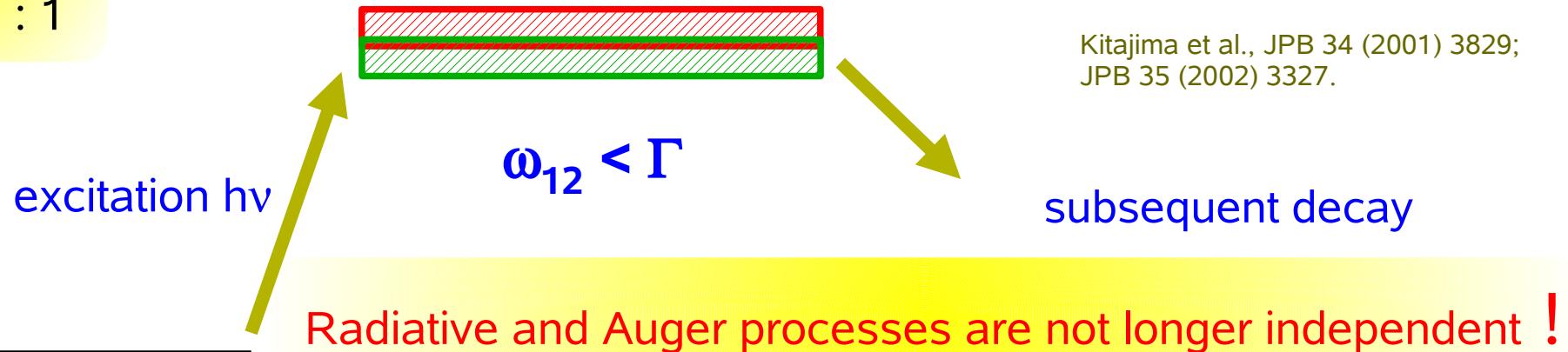


Ne:	500 : 1
Ar:	80 : 1
Kr:	25 : 1
Xe:	8 : 1

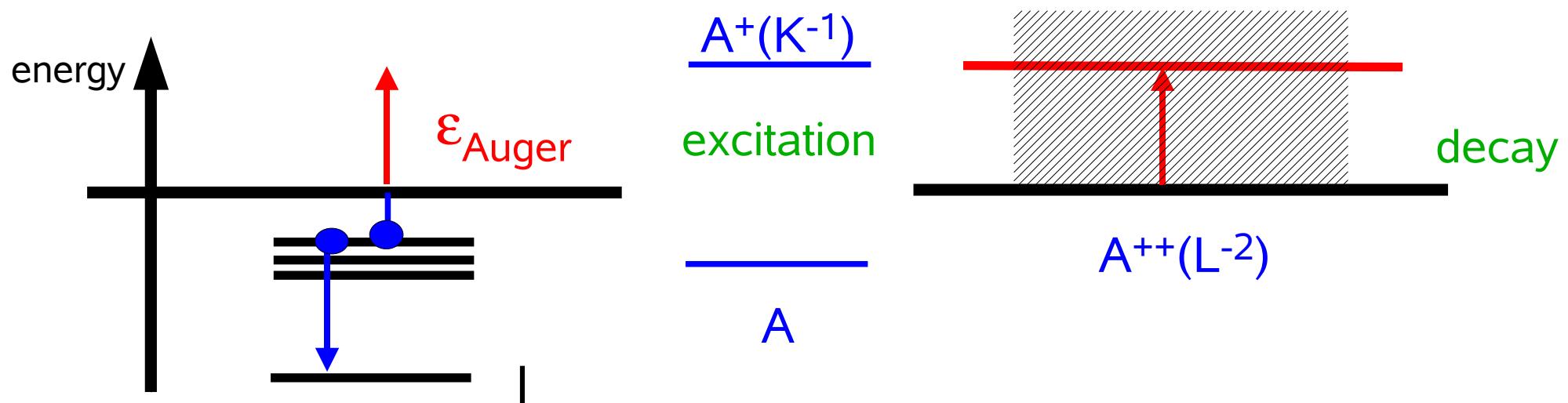
$$\frac{A_{\text{resonance}}}{A_{\text{intercombination}}}$$



Kitajima et al., JPB 34 (2001) 3829;
JPB 35 (2002) 3327.



Auger emission of excited atomic states



$$H_{DCB} = \sum_i h_D(i) + \sum_{i < j} \frac{1}{r_{ij}} + \sum_{i < j} \frac{1}{2r_{ij}} [\alpha_i \alpha_j + \frac{(\alpha_i r_i)(\alpha_j r_j)}{r_{ij}^2}]$$

Wentzel's ansatz: Autoionization is caused by electron-electron interactions which cannot be considered in an one-particle picture.

$$\sum_{i < j} \frac{1}{r_{ij}} + b(i, j) - \sum_i u(r_i)$$

Breit interaction

Summary and outlook

- ➊ Accurate atomic data are needed (more or less urgently) for a wide range of applications.
- ➋ Atomic physics still provides a great „playground“ for studying many-particle processes and electronic correlations.
- ➌ New numerical techniques have to meet the requirements for a whole „class of systems“ and not only provide 'proofs of principle'.
- ➍ Complexity of (atomic) many-particle systems: Development of ab-initio methods cannot always be separated from the processes and properties; overlap with experimental progress.
- ➎ Present and future challenges:
 - Improved treatment of open-shell structures and highly excited states
 - Coupling of bound-state densities to the continuum
 - (capture and emission of electrons, multi-photon processes,
 - Fano resonances, „complete experiments“)

