

Atomic Data for Lowly-Charged Heavy Ions

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ADAS 2013 (Bad Honnef, Germany)

Outline

- Introduction
- HFR+CPOL Method
- Results: Te II-III, Rh III, Pd III, Ag III, Pr IV
- Conclusions & Perspectives

Introduction

	IA																		VIIIA	
1	1.008 1H																			4.003 2He
2	6.941 3Li	9.012 4Be																		
3	22.990 11Na	24.305 12Mg																		
			IIIB	IVB	VB	VIB	VII B	VIII B			IB	IIB								
4	39.098 19K	40.08 20Ca	44.96 21Sc	47.88 22Ti	50.94 23V	52.00 24Cr	54.94 25Mn	55.85 26Fe	58.93 27Co	58.69 28Ni	63.546 29Cu	65.38 30Zn	69.72 31Ga	72.59 32Ge	74.92 33As	78.96 34Se	79.904 35Br	83.80 36Kr		
5	85.47 37Rb	87.62 38Sr	88.91 39Y	91.22 40Zr	92.91 41Nb	95.94 42Mo	(98) 43Tc	101.1 44Ru	102.91 45Rh	106.4 46Pd	107.87 47Ag	112.41 48Cd	114.82 49In	118.69 50Sn	121.75 51Sb	127.60 52Te	126.90 53I	131.29 54Xe		
6	132.91 55Cs	137.33 56Ba	138.91 57La	178.49 72Hf	180.95 73Ta	183.85 74W	186.2 75Re	190.2 76Os	192.2 77Ir	195.08 78Pt	196.97 79Au	200.59 80Hg	204.38 81Tl	207.2 82Pb	208.98 83Bi	(244) 84Po	(210) 85At	(222) 86Rn		
7	(223) 87Fr	226.03 88Rd	227.03 89Ac																	

Lanthanide Series

140.12 58Ce	140.9077 59Pr	144.24 60Nd	(145) 61Pm	150.36 62Sm	151.96 63Eu	157.25 64Gd	158.93 65Tb	162.50 66Dy	164.93 67Ho	167.26 68Er	168.93 69Tm	173.04 70Yb	174.97 71Lu
232.04 90Th	231.0359 91Pa	238.03 92U	237.05 93Np	(244) 94Pu	(243) 95Am	(247) 96Cm	(247) 97Bk	(251) 98Cf	(254) 99Es	(257) 100Fm	(258) 101Md	(259) 102No	(260) 103Lr

Actinide Series

Introduction

- Astrophysics: nucleosynthesis, neutron capture elements (s- and r-processes)
- Fusion: plasma facing materials, contamination, radiative loss, influx ...
- Lightening technology: richness of lanthanide spectra
- Photonics, lasers: triply-ionized lanthanides

The HFR+CPOl Method

The Relativistic Hartree-Fock (HFR) method of R.D. Cowan:
(The Theory of Atomic Structure and Spectra, Univ. of California Press, Berkeley, 1981)

Multiconfiguration approach through superpositions of configurations

Most important relativistic effects included (spin-orbit, mass-velocity correction, Darwin term, kappa-averaged orbitals)

Good agreement with fully relativistic methods

Convergence problems do occur very rarely

Can be used both in *ab initio* or semi-empirically

The HFR+CPOLE Method

The Semi-Empirical Optimization:

(R.D. Cowan, The Theory of Atomic Structure and Spectra, Univ. of California Press, Berkeley, 1981)

Radial parameters (average energies, electrostatic integrals, spin-orbit parameters) adjusted to minimize the discrepancies between the Hamiltonian eigenvalues and the experimental level energies

- Optimization of the wavefunctions
- Optimization of the wavelengths
- Optimization of the transition rates

→ Depends on the availability of experimental level energies!

The HFR+CPOL Method

The Core-Polarization Effects (HFR+CPOL):

(see e.g. Quinet et al 1999, MNRAS 307,934 & 2002, J. Alloys Comp. 344, 255)

Intravalence correlation: explicit multiconfiguration expansions

Core-valence correlation: core-polarization model potential
depending upon two parameters:

(Migdalek & Baylis 1978, J Phys B 11, L497)

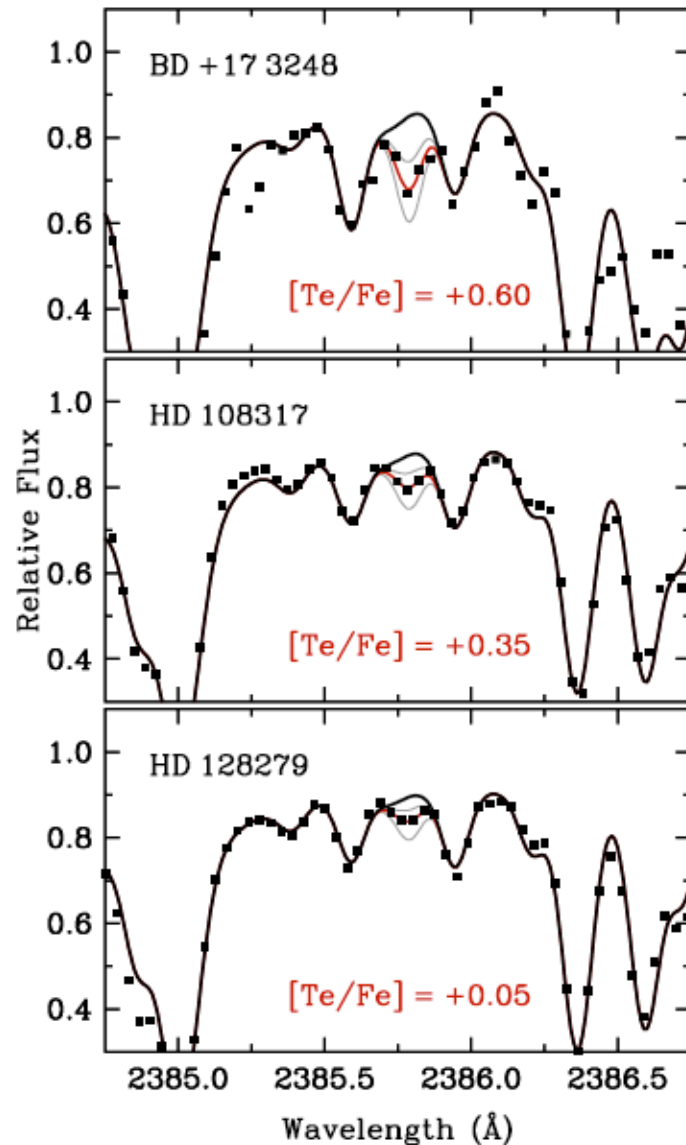
1-electric dipole polarizability of the ionic core, α_d

2-cut-off radius (size of the ionic core), r_c

Penetration of the core by valence electrons: core penetration
correction

(Hameed et al 1968, J Phys B 1, 822; Hameed 1972, J Phys B 5, 746)

Tellurium is seen in stars (Te I)



Roederer et al 2012 ApJ 747, L8:

Te I λ 2385 in HST spectra of metal-poor stars

Tellurium: no radiative data available for Te II-III

	IA																	VIII A	
1	1.008 1H																		4.003 2He
2	6.941 3Li	9.012 4Be																	
3	22.990 11Na	24.305 12Mg																	
4	39.098 19K	40.08 20Ca	44.96 21Sc	47.88 22Ti	50.94 23V	52.00 24Cr	54.94 25Mn	55.85 26Fe	58.93 27Co	58.69 28Ni	63.546 29Cu	65.38 30Zn	69.72 31Ga	72.59 32Ge	74.92 33As	78.96 34Se	79.904 35Br	83.80 36Kr	
5	85.47 37Rb	87.62 38Sr	88.91 39Y	91.22 40Zr	92.91 41Nb	95.94 42Mo	(98) 43Tc	101.1 44Ru	102.91 45Rh	106.4 46Pd	107.87 47Ag	112.41 48Cd	114.82 49In	118.69 50Sn	121.75 51Sb	127.60 52Te	126.90 53I	131.29 54Xe	
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7	(223) 87Fr	226.03 88Rd	227.03 89Ac																

Z=52 (5th Period, Group VIA)

Lanthanide Series	140.12 58Ce	140.9077 59Pr	144.24 60Nd	(145) 61Pm	150.36 62Sm	151.96 63Eu	157.25 64Gd	158.93 65Tb	162.50 66Dy	164.93 67Ho	167.26 68Er	168.93 69Tm	173.04 70Yb	174.97 71Lu
Actinide Series	232.04 90Th	231.0359 91Pa	238.03 92U	237.05 93Np	(244) 94Pu	(243) 95Am	(247) 96Cm	(247) 97Bk	(251) 98Cf	(254) 99Es	(257) 100Fm	(258) 101Md	(259) 102No	(260) 103Lr

Te II

HFR+CPOL model:

Intravalence Correlation (43 configurations):

$5p^3+5p^26p+5p^27p+5p^24f+5p^25f+5p^26f+5d^26p+5d^26f+6s^27p+$
 $5d^27p+4f^25p+4f^26p+5s5p^36s+5s5p^35d+5s5p^36d+5s5p^26p5d+$
 $5s5p^26p6d+ 5s5p^24f5d+ 5s5p^24f6d+5p^5$ (odd parity)

$5s5p^4+5p^25d+ 5p^26d+ 5p^27d+ 5p^26s+5p^27s+ 5p^28s+ 5p^25g+$
 $5p^26g+5d^25g+ 5d^26g+ 5f^25g+ 5f^26g+5s5p^36p+5s5p^34f+$
 $5s5p^35f+ 5s5p^36f+5s5p^26s5d+ 5s5p^26s6d+ 5s5p^25d6d+$
 $5s5p^26s^2+ 5s5p^25d^2$ (even parity)

Te II

HFR+CPOL model:

Core-Polarization Potential:

Pd-like Te^{6+} $[\text{Kr}]4d^{10}$ ionic core with $\alpha_d = 1.295 a_0^3$ (Fraga et al 1976, Handbook of Atomic Data, Amsterdam: Elsevier) and $r_c = \langle r \rangle_{4d} = 0.964 a_0$.

Semi-Empirical Optimization:

Odd parity: 45 experimental levels belonging to the configurations $5p^3+5p^2nl$ ($nl=6p,7p,4f$) (NIST database).
Average deviation = 87 cm^{-1} .

Even parity: 81 experimental levels belonging to the configurations $5s5p^4+5p^2nl$ ($nl=6s,7s,8s,5d,6d,7d$)(NIST database).
Average deviation = 217 cm^{-1} .

Te II

The transition probabilities and oscillator strengths for 439 strong ($\log gf > -1$) E1 transitions in the spectral range 77-997 nm.

No lifetime measurements are available for comparison!

→ f-values have been compared with an independent model to assess the reliability: the MCDHF method has been used.

Te II

Multiconfiguration Dirac-Hartree-Fock (MCDHF) model:

Fully relativistic method that takes into account the QED effects

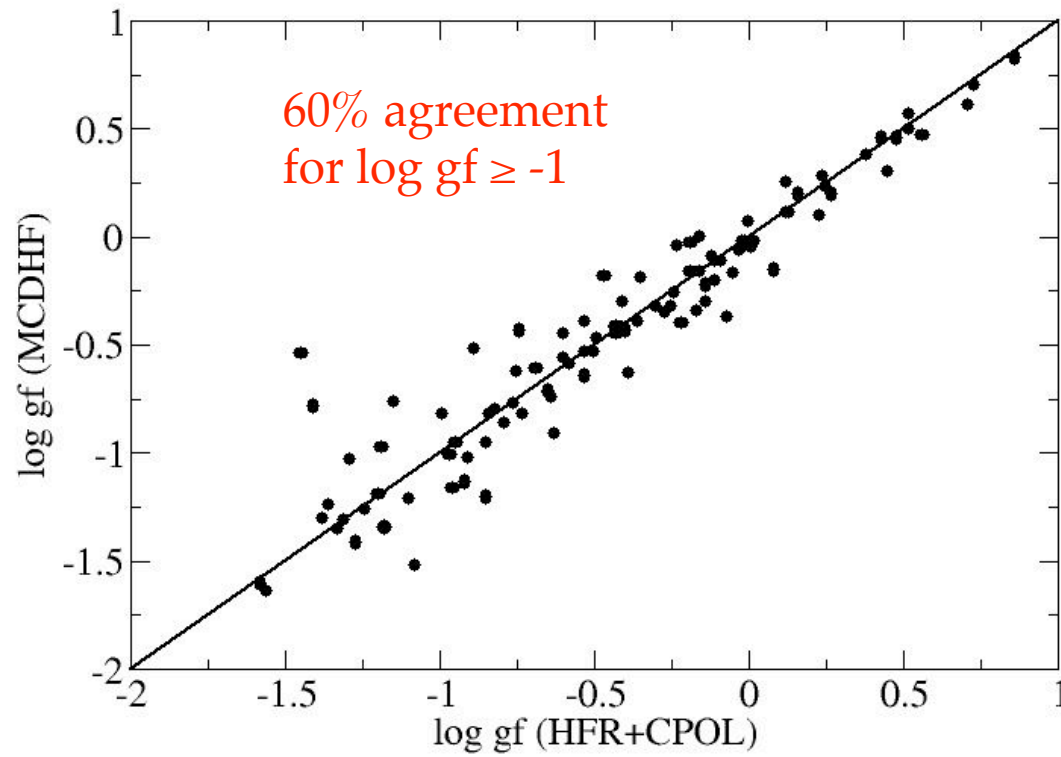
The GRASP2K package has been used (Jonsson et al 2007, CPC 177, 597)

Configuration space: generated from the multireference $5s^25p^3+5s5p^4+5s^25p^2nl$ ($nl=5d,6s,6p$) by single & double electron excitations involving the orbitals $4f,ns,np,nd$ ($n=5,6,7,8$) (102,359 CSFs).

Orbital optimization: EOL on the 70 levels of the multireference.

→ No core-valence effects due to the opening of the $n \leq 4$ shells !

Te II



Te III

HFR+CPOL model:

Intravalence Correlation (48 configurations):

$5p^2+5p6p+5p7p+5p4f+5p5f+5p6f+5d6p+5d6d+6s^2+$
 $5d^2+4f^2+5f^2+5s5p^26s+5s5p^25d+5s5p^26d+5s5p6p5d+$
 $5s5p6p6d+5s5p4f5d+5s5p4f6d+5p^4+5p^34f+5p^35f+$
 $5p^36f$ (odd parity)

$5s5p^3+5p5d+5p6d+5p7d+5p6s+5p7s+5p8s+5p5g+$
 $5p6g+5d^25g+5d6p+5d4f+5d5f+5d6f+5s5p^26p+5s5p^24f+$
 $5s5p^25f+5s5p^26f+5s5p6s5d+5s5p6s6d+5s5p5d6d+$
 $5s5p6s^2+5s5p5d^2+5p^36s+5p^35d+5p^36d$ (even parity)

Te III

HFR+CPOL model:

Core-Polarization Potential:

Pd-like Te^{6+} $[\text{Kr}]4d^{10}$ ionic core with $\alpha_d = 1.295 a_0^3$ (Fraga et al 1976, Handbook of Atomic Data, Amsterdam: Elsevier) and $r_c = \langle r \rangle_{4d} = 0.964 a_0$.

Semi-Empirical Optimization:

Even parity: 14 experimental levels belonging to the configurations $5p^2+5p6p$ (NIST database; Tauheed & Naz 2011, J. Korean Phys. Soc. 59, 2910).

Average deviation = 100 cm^{-1} .

Odd parity: 55 experimental levels belonging to the configurations $5s5p^3+5pnl$ ($nl=6s,7s,8s,5d,6d,7d$) (Tauheed & Naz 2011, J. Korean Phys. Soc. 59, 2910). Average deviation = 126 cm^{-1} .

Te III

The transition probabilities and oscillator strengths for 284 E1 transitions in the spectral range 52-901 nm.

Here again no lifetime measurements are available for comparison!

→ f-values have been also compared with a similar MCDHF model using GRASP2K.

Te III

Multiconfiguration Dirac-Hartree-Fock (MCDHF) model:

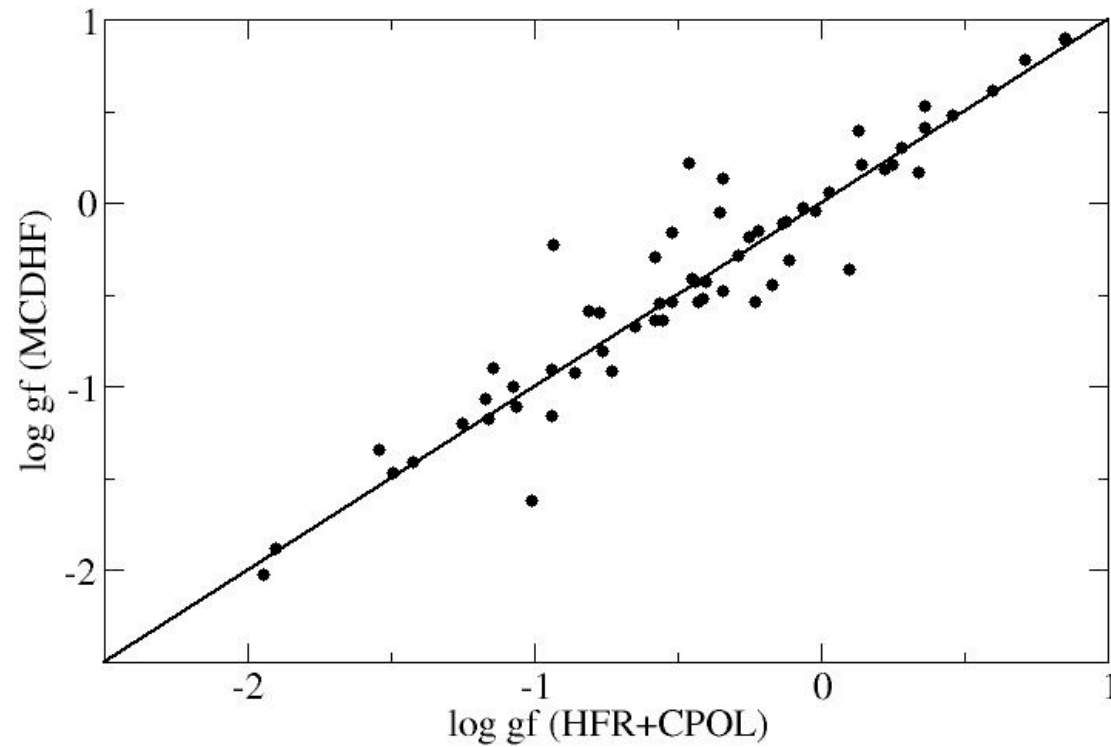
Configuration space: generated from the multireference $5s^25p^2+5s5p^3+5s^25pnl$ ($nl=5d,6s,6p$) by single & double electron excitations involving the orbitals $4f,ns,np,nd$ ($n=5,6,7,8$) (32,724 CSFs).

Orbital optimization: EOL on the 41 levels of the multireference.

→ Here also no core-valence effects due to the opening of the $n \leq 4$ shells.

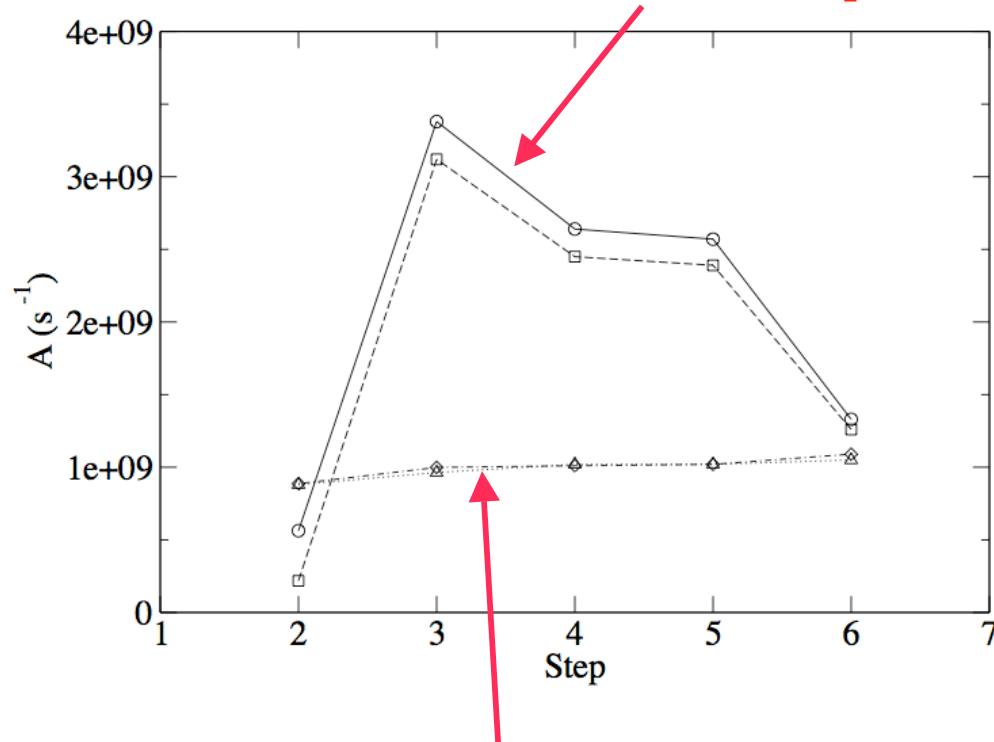
Te III

60% agreement excluding 9 transitions with convergence problem in MCDHF.



Te III: MCDHF convergence

$5p^2\ ^3P_1 - 5p6s\ ^1P^{\circ}_1$: convergence problem!
(circles: Babushkin ; squares: Coulomb)



- Step 1: ground configuration orbitals (no transition calculated)
- Step 2: multiref {5s,5p,5d,6s,6p}
- Step 3: SD {5s,5p,5d,6s,6p,6d}
- Step 4: SD {5s,5p,5d,6s,6p,6d,7s,7p,7d}
- Step 5: SD {5s,5p,5d,6s,6p,6d,7s,7p,7d,8s,8p,8d}
- Step 6: SD {5s,5p,5d,6s,6p,6d,7s,7p,7d,8s,8p,8d,4f}

$5p^2\ ^3P_2 - 5p6s\ ^3P^{\circ}_1$: converged!
(diamonds: Babushkin ; triangles: Coulomb)

Rh III, Pd III & Ag III: no radiative rates available!

Z=45-47 (5th Period, Groups VIII B-IB)

	IA																		VIIIA	
1	1.008 1H																			4.003 2He
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			IIIB	IVB	VB	VIB	VII B	VIII B			IB	IIB	IIIA	IVA	VA	VIA	VIIA			
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Rh III, Pd III & Ag III: Computational Strategy

- Rh III: similar model as in the isoelectronic ion Ru II (Palmeri et al 2009, J Phys B 42, 165005) → good agreement with the TR-LIF lifetime measurements (within a few percents).
 - Pd III: similar model as in Rh II (Quinet et al 2012, A&A 537, A74) → also good agreement with the TR-LIF lifetime measurements (within a few percents).
 - Ag III: similar model as in Pd III adding one electron to the 4d subshell
- An accuracy of a few percents for the radiative rates is expected at least for the strongest lines

Rh III, Pd III & Ag III: HFR+CPOL Models (CI)

- Rh III: $4d^7 + 4d^65s + 4d^66s + 4d^65d + 4d^66d + 4d^55s^2 + 4d^55p^2 + 4d^55d^2 + 4d^55s6s$ (even parity) & $4d^65p + 4d^66p + 4d^64f + 4d^65f + 4d^55s5p + 4d^55s6p + 4d^55p6s$ (odd parity); CPOL: Rh V $4d^5$ core with $\alpha_d=3.31 a_0^3$ (Fraga et al, 1976) & $r_c=\langle r \rangle_{4d}=1.43 a_0$; Fit:
- Pd III: $4d^8 + 4d^75s + 4d^76s + 4d^75d + 4d^76d + 4d^65s^2 + 4d^65p^2 + 4d^65d^2 + 4d^65s6s + 4d^65s5d + 4d^65s6d$ (even parity) & $4d^75p + 4d^76p + 4d^74f + 4d^75f + 4d^65s5p + 4d^65s6p + 4d^65p5d + 4d^65p6s$ (odd parity)
- Ag III: $4d^9 + 4d^85s + 4d^86s + 4d^85d + 4d^86d + 4d^75s^2 + 4d^75p^2 + 4d^75d^2 + 4d^75s6s + 4d^75s5d + 4d^75s6d$ (even parity) & $4d^85p + 4d^86p + 4d^84f + 4d^85f + 4d^75s5p + 4d^75s6p + 4d^75p5d + 4d^75p6s$ (odd parity)

Rh III, Pd III & Ag III: HFR+CPOL Models (CPOL)

- Rh III: Rh V $4d^5$ core with $\alpha_d=3.31 a_0^3$ (Fraga et al, 1976) & $r_c=\langle r \rangle_{4d}=1.43 a_0$
- Pd III: Pd V $4d^6$ core with $\alpha_d=3.17 a_0^3$ (Fraga et al, 1976) & $r_c=\langle r \rangle_{4d}=1.36 a_0$
- Ag III: Ag V $4d^7$ core with $\alpha_d=3.04 a_0^3$ (Fraga et al, 1976) & $r_c=\langle r \rangle_{4d}=1.04 a_0$

Rh III, Pd III & Ag III: HFR+CPOLE Models (Fits)

- Rh III: 196 experimental levels belonging to $4d^7$, $4d^65s$ & $4d^65p$ (NIST database). Average deviations: 28 cm^{-1} (even parity) & 84 cm^{-1} (odd parity).
- Pd III: 177 experimental levels belonging to $4d^8$, $4d^75s$, $4d^76s$, $4d^75p$ & $4d^65s5p$ (NIST database). Average deviations: 80 cm^{-1} (even parity) & 68 cm^{-1} (odd parity).
- Ag III: 64 experimental levels belonging to $4d^9$, $4d^85s$ & $4d^85p$ (NIST database). Average deviations: 117 cm^{-1} (even parity) & 66 cm^{-1} (odd parity).

Experimental and Calculated Even-Parity Energy Levels in Ag III

E_{exp}^a (cm^{-1})	E_{calc}^b (cm^{-1})	ΔE (cm^{-1})	J	LS composition ^c (%)
Even parity				
0.0	0059	-59	5/2	99 $4d^9 \ ^2D$
4609.2	4550	59	3/2	99 $4d^9 \ ^2D$
63246.3	63269	-23	9/2	99 $4d^8(^3F)5s \ ^4F$
65759.2	65714	46	7/2	90 $4d^8(^3F)5s \ ^4F$ + 9 $4d^8(^3F)5s \ ^2F$
68139.1	68050	89	5/2	95 $4d^8(^3F)5s \ ^4F$
69345.7	69268	77	3/2	91 $4d^8(^3F)5s \ ^4F$ + 8 $4d^8(^1D)5s \ ^2D$
71686.4	71715	-29	7/2	89 $4d^8(^3F)5s \ ^2F$ + 9 $4d^8(^3F)5s \ ^4F$
73928.9	74127	-198	5/2	46 $4d^8(^3F)5s \ ^2F$ + 27 $4d^8(^3P)5s \ ^4P$ + 26 $4d^8(^1D)5s \ ^2D$
76402.3	76480	-78	5/2	54 $4d^8(^3P)5s \ ^4P$ + 42 $4d^8(^3F)5s \ ^2F$
77408.7	77555	-147	3/2	65 $4d^8(^3P)5s \ ^4P$ + 22 $4d^8(^1D)5s \ ^2D$ + 8 $4d^8(^3P)5s \ ^2P$
78893.2	78959	-66	1/2	98 $4d^8(^3P)5s \ ^4P$
80127.4	80279	-151	3/2	47 $4d^8(^1D)5s \ ^2D$ + 34 $4d^8(^3P)5s \ ^4P$ + 14 $4d^8(^3P)5s \ ^2P$
82228.4	82421	-193	5/2	70 $4d^8(^1D)5s \ ^2D$ + 17 $4d^8(^3P)5s \ ^4P$ + 10 $4d^8(^3F)5s \ ^2F$
85179.8	85157	23	3/2	76 $4d^8(^3P)5s \ ^2P$ + 22 $4d^8(^1D)5s \ ^2D$
85505.5	85421	84	1/2	97 $4d^8(^3P)5s \ ^2P$
85596.1	85293	303	9/2	98 $4d^8(^1G)5s \ ^2G$
85724.1	85351	373	7/2	98 $4d^8(^1G)5s \ ^2G$
111436.4	111551	-115	1/2	96 $4d^8(^1S)5s \ ^2S$

Rh III, Pd III & Ag III: Radiative Rates

- Rh III: 2150 E1 transitions in the spectral region 76-519 nm
- Pd III: 2120 E1 transitions in the spectral region 61-975 nm
- Ag III: 440 E1 transitions in the spectral region 59-801 nm

Sample of E1 Radiative Rates ($\log gf > -0.5$) in Rh III

Wavelength ^a (nm)	Lower level ^b			Upper level ^b			$\log gf^c$	gA^c (s ⁻¹)
	E (cm ⁻¹)	Parity	J	E (cm ⁻¹)	Parity	J		
85.228	2148	(e)	7/2	119481	(o)	7/2	-0.48	3.01E+09
85.477	0	(e)	9/2	116991	(o)	9/2	-0.36	3.98E+09
85.988	0	(e)	9/2	116296	(o)	9/2	0.02	9.55E+09
86.133	3486	(e)	5/2	119586	(o)	5/2	-0.34	4.15E+09
86.200	15130	(e)	7/2	131138	(o)	5/2	-0.42	3.40E+09
86.211	3486	(e)	5/2	119481	(o)	7/2	-0.28	4.73E+09
86.376	2148	(e)	7/2	117921	(o)	7/2	-0.15	6.29E+09
86.622	0	(e)	9/2	115445	(o)	11/2	-0.47	3.04E+09
86.758	4322	(e)	3/2	119586	(o)	5/2	-0.44	3.18E+09
87.038	13030	(e)	9/2	127923	(o)	7/2	0.27	1.64E+10
87.076	2148	(e)	7/2	116991	(o)	9/2	-0.11	6.85E+09
87.425	3486	(e)	5/2	117870	(o)	5/2	-0.44	3.14E+09
87.786	13030	(e)	9/2	126944	(o)	9/2	-0.20	5.51E+09
87.839	2148	(e)	7/2	115992	(o)	5/2	-0.33	4.06E+09
89.432	0	(e)	9/2	111817	(o)	9/2	-0.25	4.65E+09
89.658	19576	(e)	9/2	131111	(o)	7/2	0.03	8.84E+09
89.755	0	(e)	9/2	111414	(o)	11/2	0.06	9.55E+09
89.765	15130	(e)	7/2	126532	(o)	7/2	-0.08	6.89E+09

Praseodymium: radiative parameters are sparse in Pr IV

Z=59 (Lanthanide Series)

	IA																		VIII A					
1	1.008 1H																	4.003 2He						
2	6.941 3Li	9.012 4Be																	10.81 5B	12.011 6C	14.007 7N	15.999 8O	18.998 9F	20.179 10Ne
3	22.990 11Na	24.305 12Mg	IIIB		IVB	VB	VIB	VII B	VIII B			IB	IIB	26.98 13Al	28.09 14Si	30.974 15P	32.06 16S	35.453 17Cl	39.948 18Ar					
4	39.098 19K	40.08 20Ca	44.96 21Sc	47.88 22Ti	50.94 23V	52.00 24Cr	54.94 25Mn	55.85 26Fe	58.93 27Co	58.69 28Ni	63.546 29Cu	65.38 30Zn	69.72 31Ga	72.59 32Ge	74.92 33As	78.96 34Se	79.904 35Br	83.80 36Kr						
5	85.47 37Rb	87.62 38Sr	88.91 39Y	91.22 40Zr	92.91 41Nb	95.94 42Mo	(98) 43Tc	101.1 44Ru	102.91 45Rh	106.4 46Pd	107.87 47Ag	112.41 48Cd	114.82 49In	118.69 50Sn	121.75 51Sb	127.60 52Te	126.90 53I	131.29 54Xe						
6	132.91 55Cs	137.33 56Ba	138.91 57La	178.49 72Hf	180.95 73Ta	183.85 74W	186.2 75Re	190.2 76Os	192.2 77Ir	195.08 78Pt	196.97 79Au	200.59 80Hg	204.38 81Tl	207.2 82Pb	208.98 83Bi	(244) 84Po	(210) 85At	(222) 86Rn						
7	(223) 87Fr	226.03 88Rd	227.03 89Ac																					

Lanthanide Series	140.12 58Ce	140.9077 59Pr	144.24 60Nd	(145) 61Pm	150.36 62Sm	151.96 63Eu	157.25 64Gd	158.93 65Tb	162.50 66Dy	164.93 67Ho	167.26 68Er	168.93 69Tm	173.04 70Yb	174.97 71Lu
Actinide Series	232.04 90Th	231.03588 91Pa	238.03 92U	237.05 93Np	(244) 94Pu	(243) 95Am	(247) 96Cm	(247) 97Bk	(251) 98Cf	(254) 99Es	(257) 100Fm	(258) 101Md	(259) 102No	(260) 103Lr

Pr IV: HFR+CPOL Model

Similar model as in the isoelectronic ion Ce III (Biémont et al 2002, MNRAS 336, 1155) which gave a good agreement (within 5%) with the TR-LIF lifetime measurements.

→ Accuracy of ~5% is expected in Pr IV

Intravalence Correlation:

$4f^2 + 4fn_p (n=6-7) + 5d^2 + 5dn_s (n=6-7) + 6s^2 + 5d6d + 4fn_f (n=5-7) + 6p^2$ (even parity)

$4fnd (n=5-7) + 4fns (n=6-8) + 5d6p + 4fng (n=5-6) + 6s6p$ (odd parity)

Pr IV: HFR+CPOL Model

Core-Polarization Potential:

Xe-like Pr⁵⁺ 5p⁶ ionic core: $\alpha_d = 5.40 a_0^3$ (Fraga et al, 1976)
 $r_c = \langle r \rangle_{5p} = 1.60 a_0$

Semi-Empirical Optimization:

Even parity: 50 4f²+5d²+4f6p+4f5f experimental levels (NIST database). Average deviation = 34 cm⁻¹.

Odd parity: 36 4f²+5d²+4f6p+4f5f experimental levels (NIST database). Average deviation = 95 cm⁻¹.

Pr IV: Radiative Rates

The transition probabilities and oscillator strengths have been calculated for :

- 199 strong ($\log gf \geq -2$) E1 transitions in the spectral region 118 - 302 nm.
- 30 forbidden (M1+E2) transitions ($gA > 0.01 \text{ s}^{-1}$) within the $4f^2$ configuration in the spectral region 450 - 22,800 nm.

Pr IV: Forbidden transitions within $4f^2$

Air wavelength ^a (nm)	Transition	Type ^b	gA (This work) (s ⁻¹)	gA (Other) ^c (s ⁻¹)
450.0902	$^3H_4 - ^1I_6$	E2	1.63E-02	2.11E-02
498.3791	$^3H_5 - ^1I_6$	M1	6.42E+00	6.29E+00
550.3866	$^3F_2 - ^3P_2$	M1+E2	6.84E-02	6.49E-02
560.9344	$^3H_6 - ^1I_6$	M1	6.31E+00	6.14E+00
587.6972	$^3F_2 - ^3P_1$	M1+F2	1.82E-01	9.11E-02
597.0146	$^3F_2 - ^3P_2$	M1+E2	8.31E-01	8.00E-01
609.8402	$^3F_2 - ^3P_0$	E2	1.10E-01	2.78E-03
613.1068	$^3F_4 - ^3P_2$	E2	1.61E-01	3.06E-02
641.1683	$^3F_3 - ^3P_1$	E2	1.67E-01	1.94E-02
650.9979	$^3F_4 - ^1I_6$	E2	2.50E-02	2.87E-02
755.1150	$^1G_4 - ^3P_2$	E2	8.68E-02	1.67E-02
810.2958	$^3F_2 - ^1D_2$	M1	3.82E+00	3.69E+00
813.4261	$^1G_4 - ^1I_6$	E2	2.28E-02	2.57E-02
915.5709	$^3F_3 - ^1D_2$	M1	5.01E+00	4.77E+00
953.9696	$^3F_4 - ^1D_2$	E2	2.31E-02	4.73E-03
1007.6623	$^3H_4 - ^1G_4$	M1	1.56E+00	1.51E+00
1286.7901	$^3H_5 - ^1G_4$	M1	1.80E+00	1.76E+00
1458.4438	$^3H_4 - ^3F_4$	M1	1.39E+00	1.37E+00
1558.3623	$^3H_4 - ^3F_3$	M1	4.27E-02	4.39E-02
1715.9100	$^1D_2 - ^3P_2$	M1	3.09E+00	
2125.8758	$^3H_5 - ^3F_4$	M1	4.49E-01	
2139.3369	$^1D_2 - ^3P_1$	M1	5.43E-01	
2851.4754	$^3F_3 - ^1G_4$	M1	2.70E+00	
3260.1682	$^3F_4 - ^1G_4$	M1	1.97E+00	
4469.0540	$^3H_5 - ^3H_6$	M1	3.22E+00	
4645.3791	$^3H_4 - ^3H_5$	M1	2.72E+00	2.60E+00
7047.1330	$^3F_2 - ^3F_3$	M1	2.72E+00	
8669.5351	$^3P_1 - ^3P_2$	M1	1.16E-01	
16185.9865	$^3P_0 - ^3P_1$	M1	1.40E-02	
22746.4098	$^3F_3 - ^3F_4$	M1	1.36E-02	

^a Deduced from experimental levels compiled by Martin *et al.* (1978)

^b Contributions larger than 1%

^c Dodson and Zia (2012)

Conclusions & Perspectives

- New radiative data have been calculated in Te II-III, Rh III, Pd III, Ag III & Pr IV using the HFR+CPOL method
- Accuracy of ~60% has been estimated through comparisons with independent MCDHF calculations in Te II-III
- Use of similar experimentally benchmarked models in isoelectronic ions for Rh III, Pd III, Ag III & Pr IV: expected accuracy of a few %
- Lifetime & branching fraction measurements are needed!
- Publications: Zhang et al (2013, A&A 551, A136); Zhang et al (2013, Phys Scr, to be published); Enzonga Yoca & Quinet (2013, in preparation)

Collaborations

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Thank you for your attention!