

Determination of divertor temperature and density using N II line ratios

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Talk Overview

1) Introduce spectroscopic techniques used in analysis

- *Balmer and impurity spectroscopy*

2) Illustrate AUG divertor setup

- *Geometric layout*
- *SOL Plasma Simulation*

3) Atomic modelling of N II emission

- *Line intensities and ratios*
- *ADAS feature generation (AFG)*
- *Framework for Feature Synthesis (FFS)*

4) Temporal behaviour of fitted parameters

- $n_{N1\text{+}}, T_e$ & n_e

Introduction – Balmer spectroscopy

Balmer series emission can provide valuable information about the divertor plasma conditions

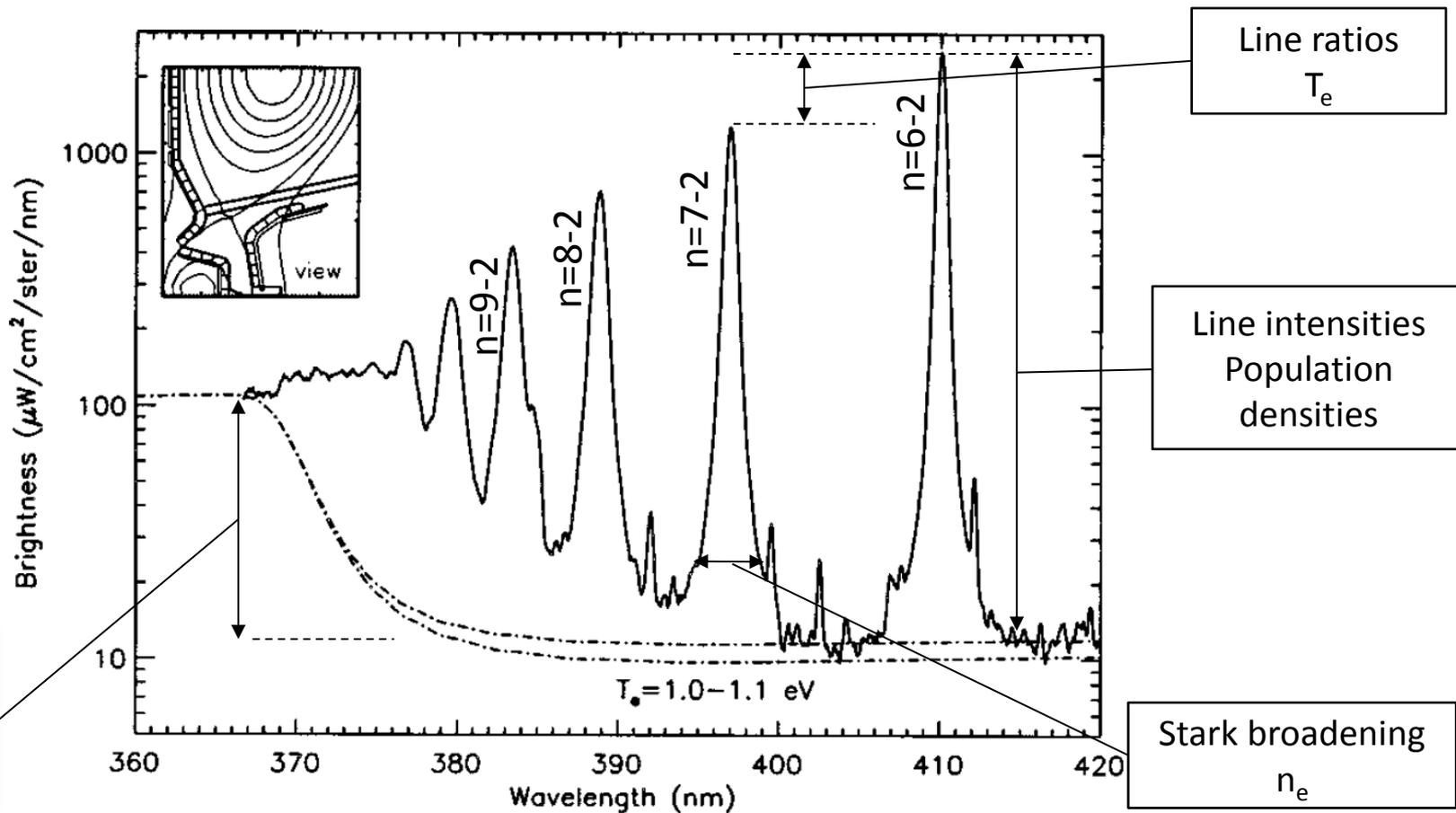


Figure 1: J.L. Terry et al. PoP 5, 1759 (1998)

Introduction – Balmer spectroscopy

A narrow spectral range is used on AUG to measure the broadening of the H_δ and H_ϵ lines

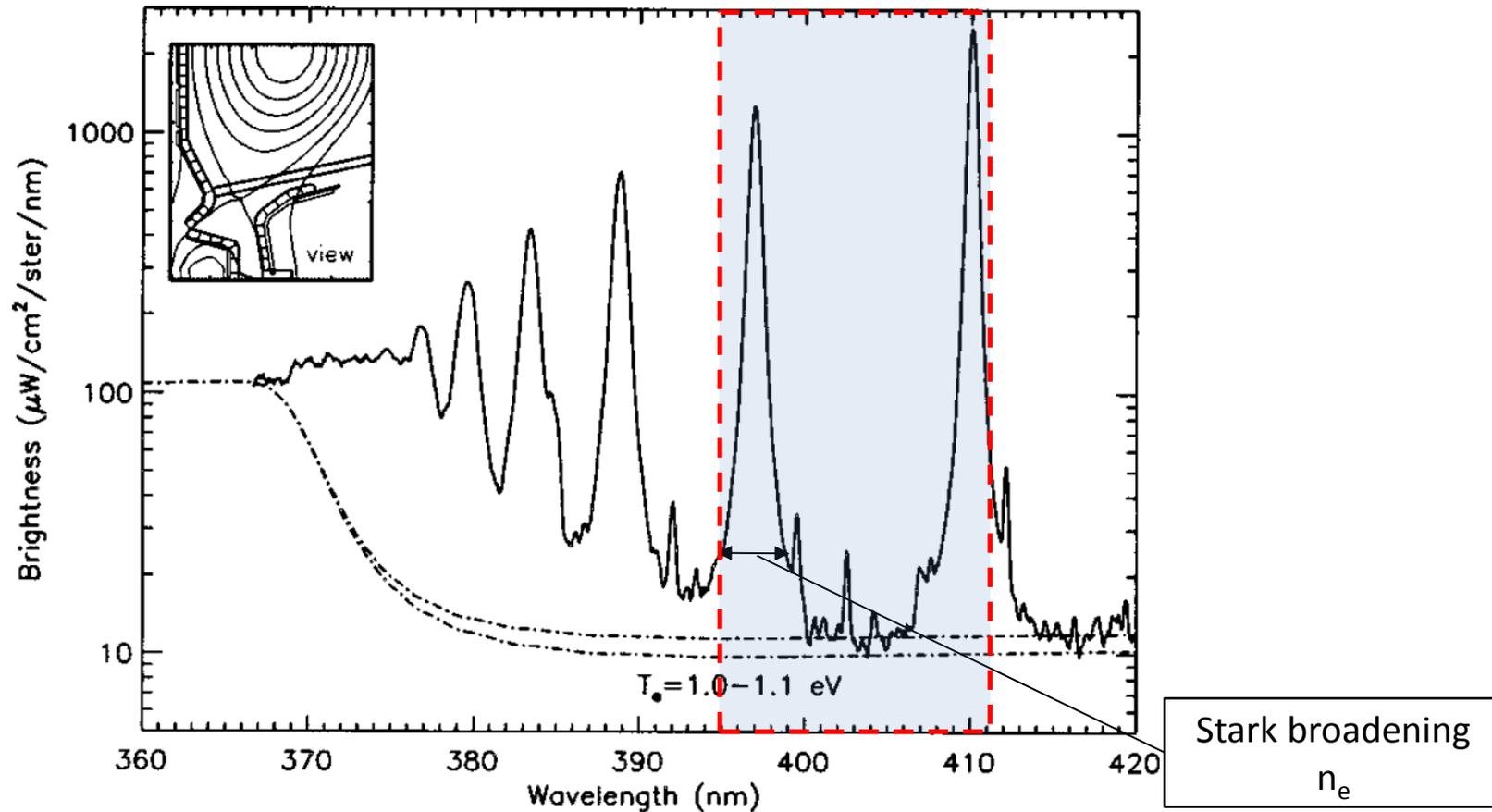


Figure 1: J.L. Terry et al. PoP 5, 1759 (1998)

Introduction – Impurity spectroscopy

A number of impurity lines (N II, N III, N IV, He I, & W I) also emit within this narrow spectral range

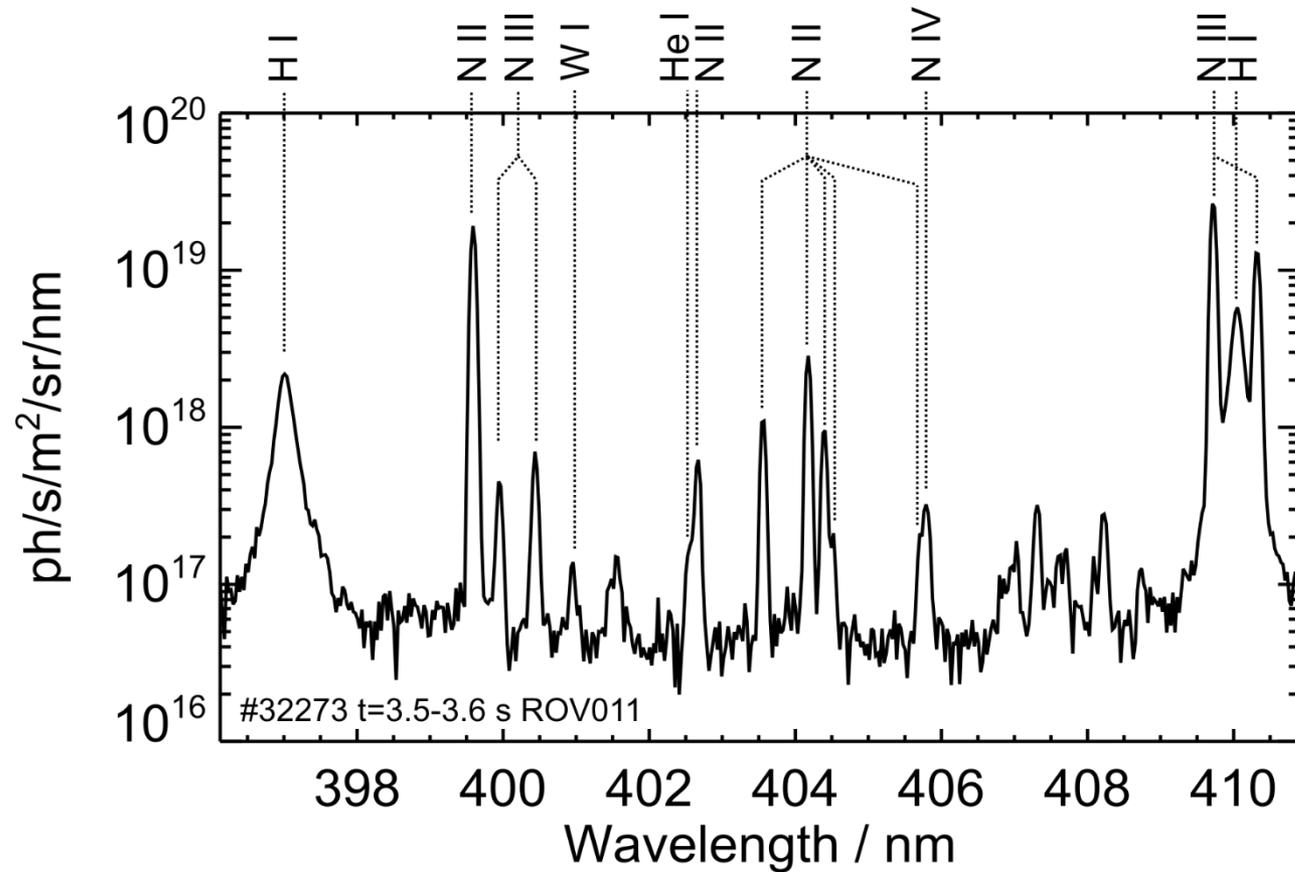


Figure 2: Spectra measured along sightline ROV011 during AUG discharge #32273

Introduction – Temperature range

The ratio of lines emanating from N II, N III, & N IV provide a measurement of n_e in hotter regions of the divertor plasma ($T_e < 20$ eV) in comparison to Balmer spectroscopy

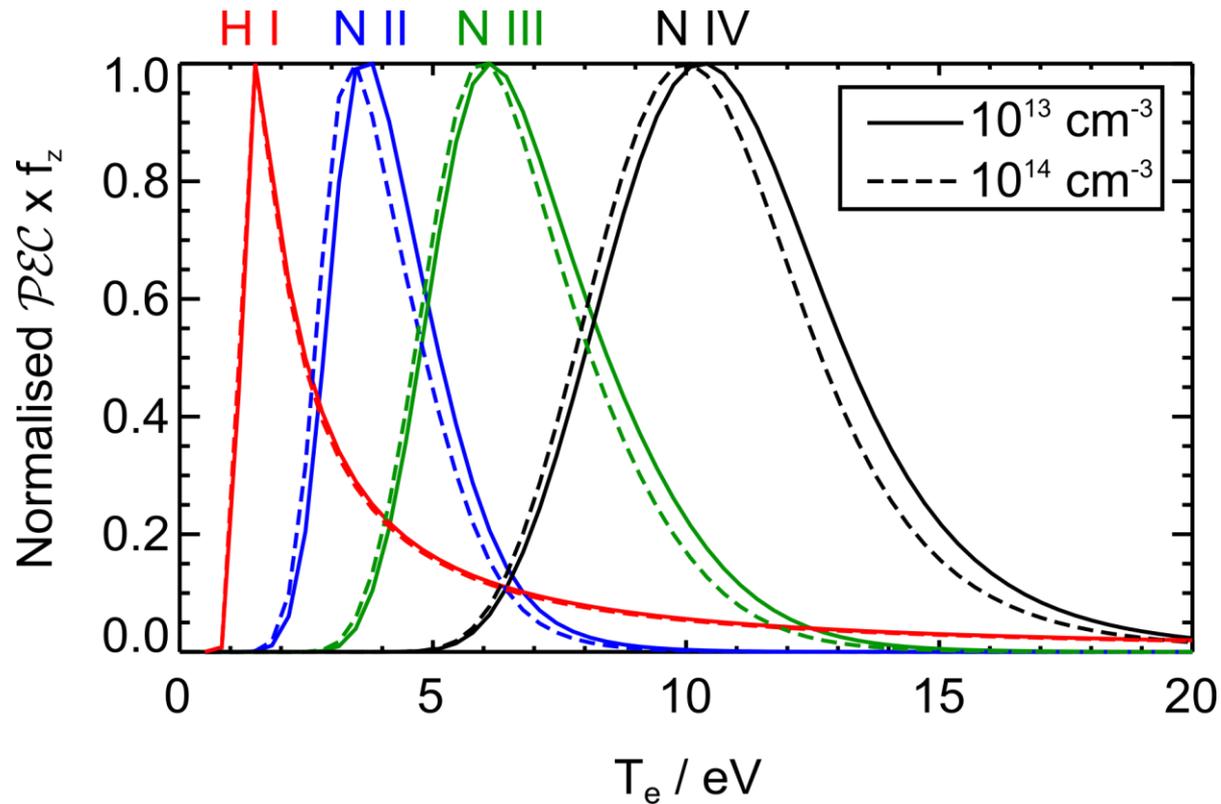
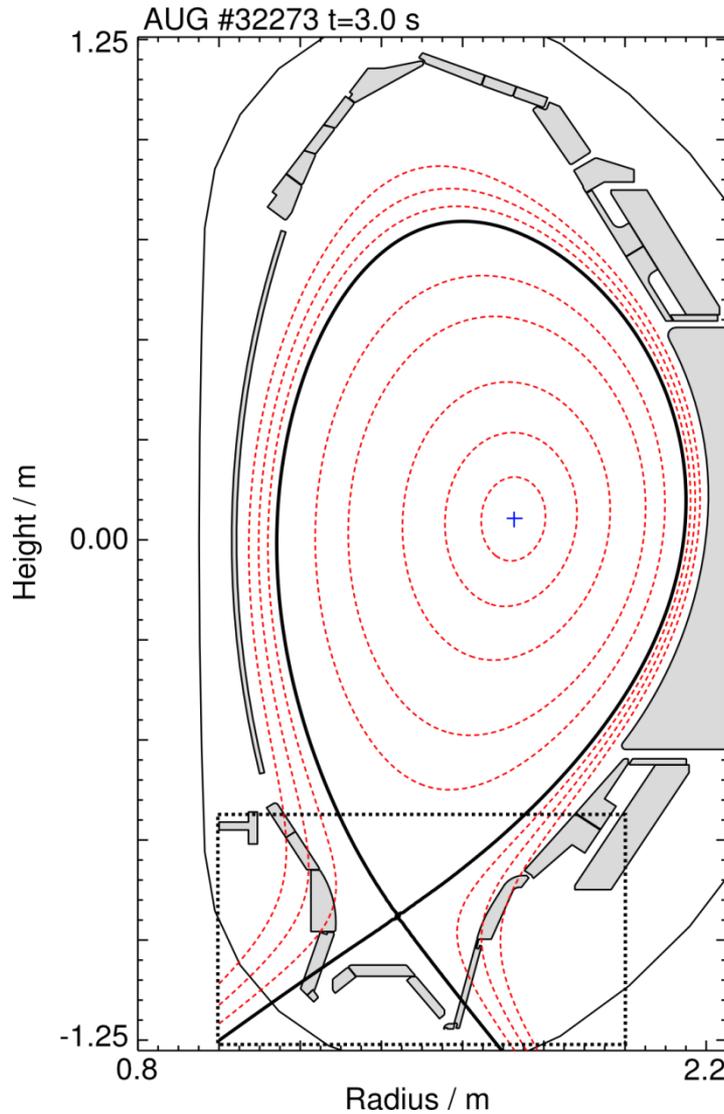


Figure 3: PECs and ionisation balance calculated using ADAS 96 data

AUG experimental setup

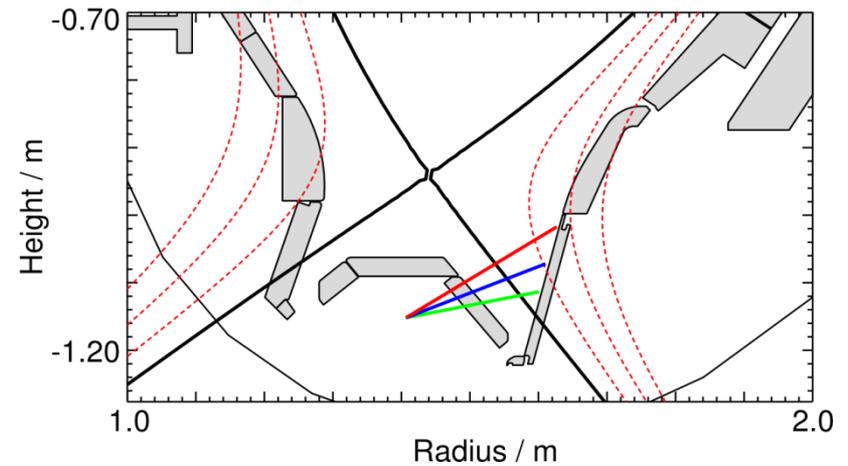


Divertor visible spectroscopy

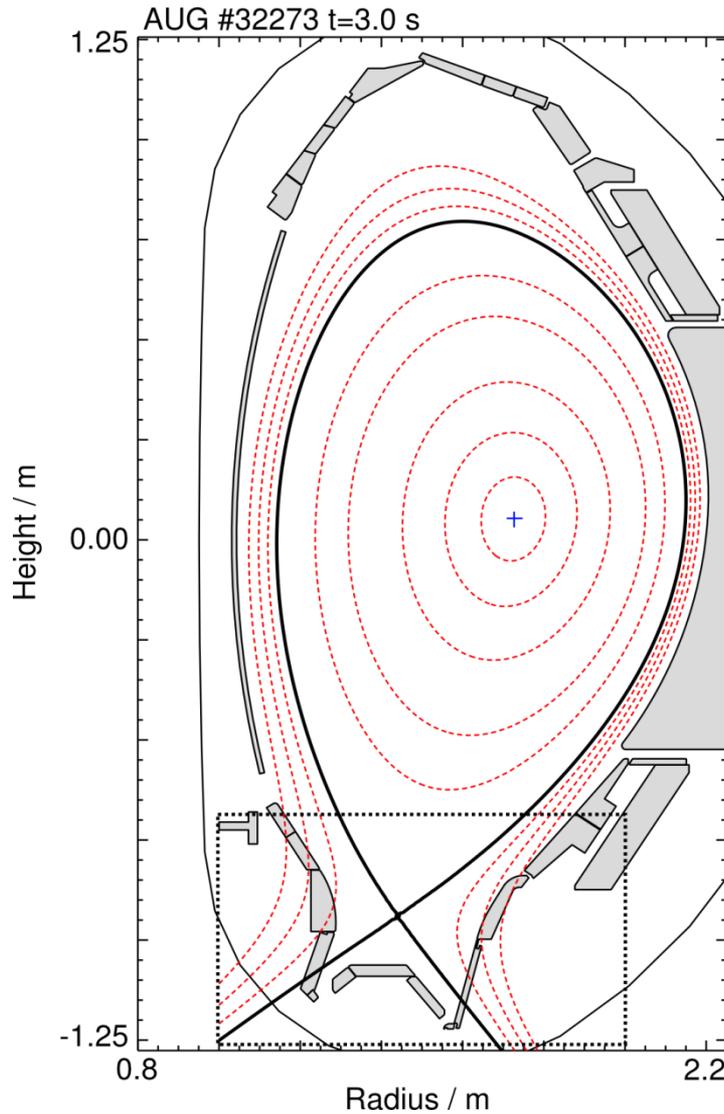
- $396 \text{ nm} < \lambda < 411 \text{ nm}$
- Instrument function: 0.08 nm
- Integration time: 2.5 ms

Sightlines

- ROV009
- ROV011
- ROV013



AUG experimental setup

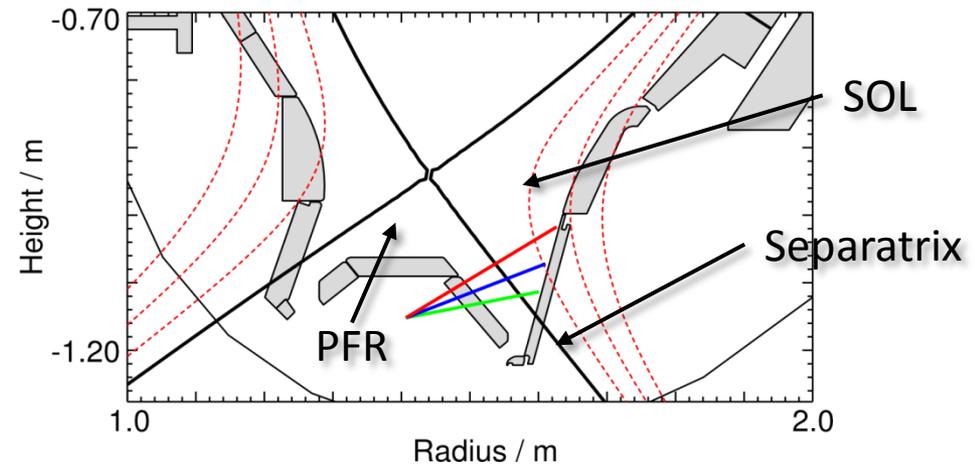


Divertor visible spectroscopy

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SOL Plasma Simulation

- SOLPS simulation of T_e and n_e
 - T_e in SOL too high (coupled with low n_e) for low charge nitrogen emission
 - N II, N III, & N IV emit locally in the PFR of the outer divertor leg

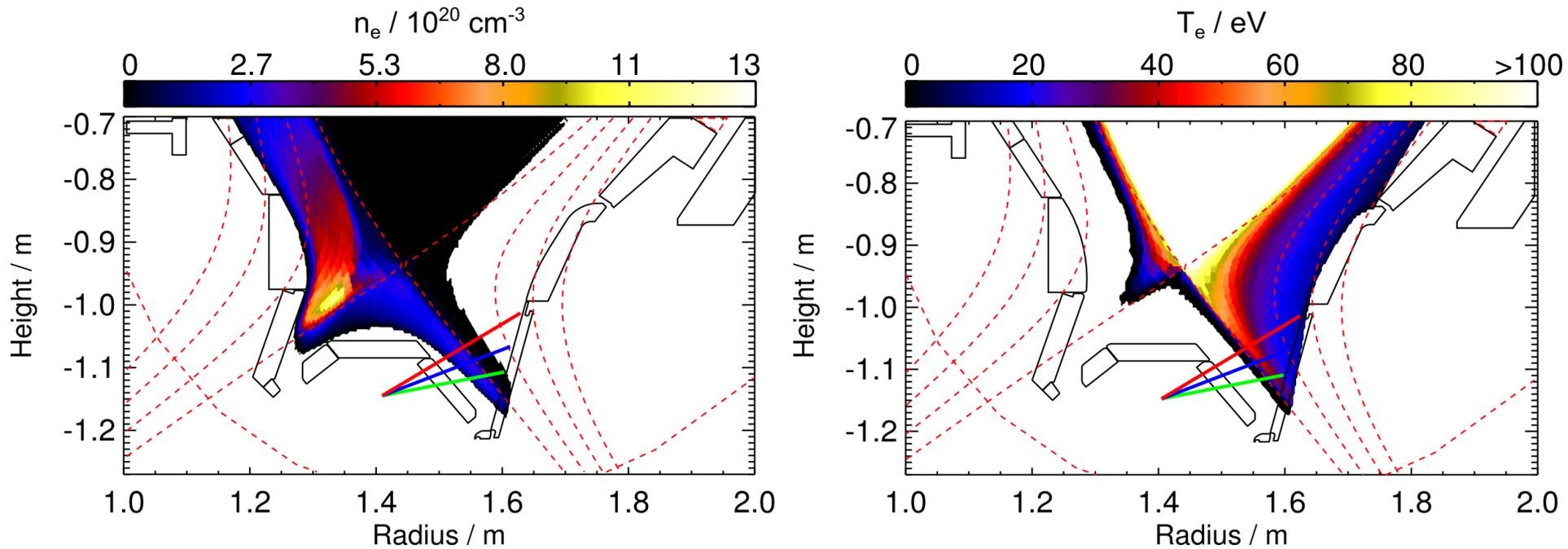


Figure 4: SOLPS simulation provided by F. Reimold

SOL Plasma Simulation

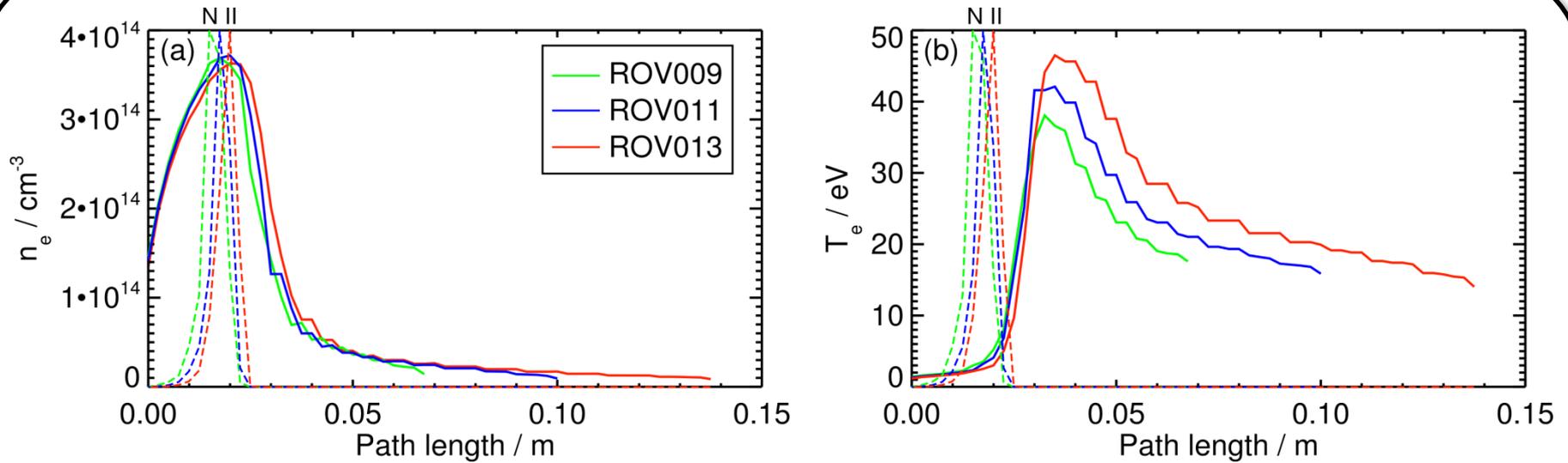


Figure 5: Sightline profiles of (a) n_e and (b) T_e as a function of the path length. Normalised N II emission is shown by the dashed lines to indicate the emitting regions.

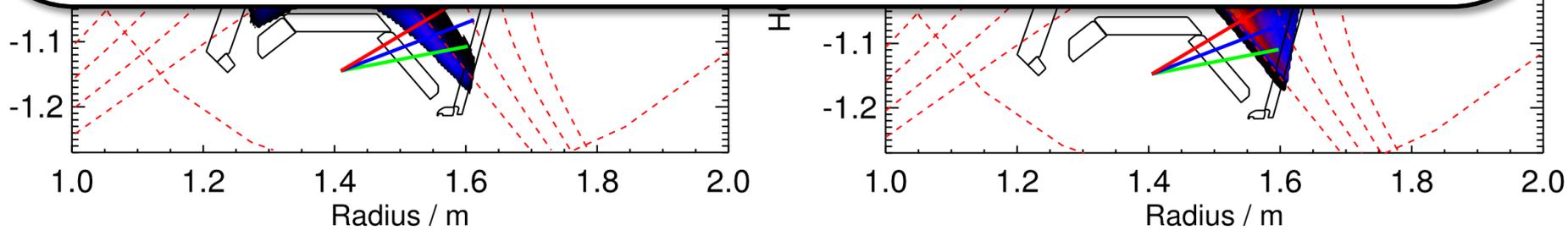


Figure 4: SOLPS simulation provided by F. Reimold

Nitrogen lines in the near UV – visible (term resolved)

(1) N II at $\lambda \sim 344$ nm: $2s^2 2p 3p \ ^1S \rightarrow 2s^2 2p 3s \ ^1P$

(2) N IV at $\lambda \sim 348$ nm: $1s^2 2s 3p \ ^3P \rightarrow 1s^2 2s 3s \ ^3S$

(3) N III at $\lambda \sim 377$ nm: $2s 2p 3p \ ^4S \rightarrow 2s 2p 3s \ ^4P$

(4) N II at $\lambda \sim 383$ nm: $2s^2 2p 4s \ ^3P \rightarrow 2s^2 2p 3p \ ^3P$

(5) N II at $\lambda \sim 395$ nm: $2s^2 2p 3p \ ^1D \rightarrow 2s^2 2p 3s \ ^3P$

(6) **N II at $\lambda \sim 399.5$ nm: $2s^2 2p 3p \ ^1D \rightarrow 2s^2 2p 3s \ ^1P$**

(7) N III at $\lambda \sim 400$ nm: $2s^2 5f \ ^2F \rightarrow 2s^2 4d \ ^2D$

(8) **N II at $\lambda \sim 402.5$ nm: $2s^2 2p 4f \ ^1G \rightarrow 2s^2 2p 3d \ ^3F$**

(9) **N II at $\lambda \sim 404.0$ nm: $2s^2 2p 4f \ ^3G \rightarrow 2s^2 2p 3d \ ^3F$**

(10) N IV at $\lambda \sim 406$ nm: $1s^2 2s 3d \ ^1D \rightarrow 1s^2 2s 3p \ ^1P$

(11) N III at $\lambda \sim 409$ nm: $2s^2 3p \ ^2P \rightarrow 2s^2 3s \ ^2S$

(12) N II at $\lambda \sim 413$ nm: $2s 2p^2 3p \ ^5S \rightarrow 2s 2p^2 3s \ ^5P$

(13) N II at $\lambda \sim 422$ nm: $2s^2 2p 4s \ ^1P \rightarrow 2s^2 2p 3p \ ^1D$

(14) N III at $\lambda \sim 420$ nm: $2s 2p 3p \ ^2D \rightarrow 2s 2p 3s \ ^2P$

- This analysis focuses on **three** N II lines
 - $\lambda \sim 399.5$ nm: $2s^2 2p 3p \ ^1D - 2s^2 2p 3s \ ^1P$
 - $\lambda \sim 402.5$ nm: $2s^2 2p 4f \ ^1G - 2s^2 2p 3d \ ^3F$
 - $\lambda \sim 404.0$ nm: $2s^2 2p 4f \ ^3G - 2s^2 2p 3d \ ^3F$
- A wider spectral range ($340 \text{ nm} < \lambda < 430 \text{ nm}$) would also facilitate an N III and N IV line ratio analysis

N II term resolved line intensities

- A new N¹⁺ ADF04 file with a more recent set of R-Matrix EIE rate coefficients (4d and 4f shell transitions supplemented by ADAS DW calculations) has been generated for this analysis
- PECs calculated with the new ADF04 file show a significant decrease in both 399.5 nm and 402.5 nm term resolved intensities in comparison to the ADAS 96 data
- Recombination does not contribute to 399.5 nm line, but has modest contribution to both 404.0 and 402.5 nm lines

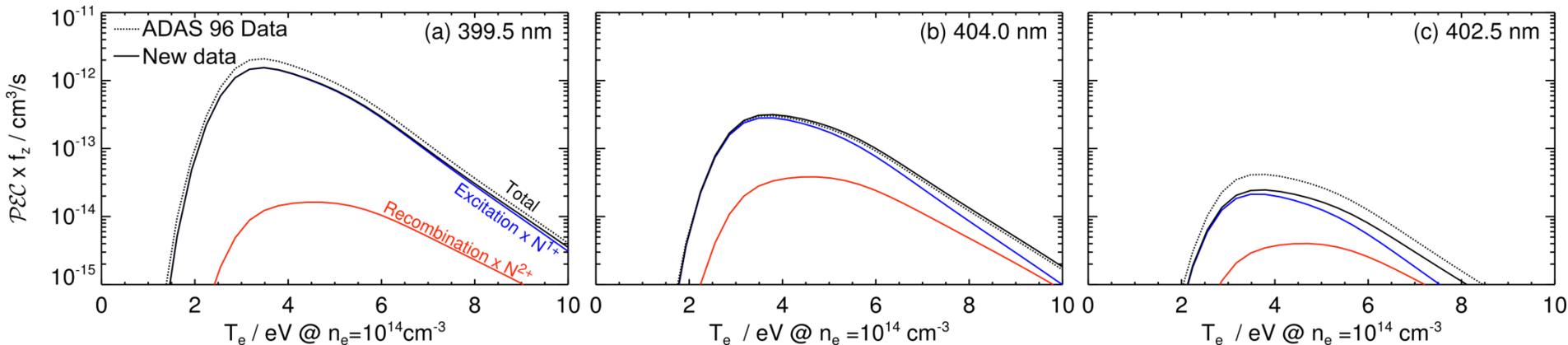


Figure 6: ADAS 96 data calculated using collision strengths from Stafford et al. *Mon. Not. R. Astr. Soc.* **268**, 816 (1994) & R. Frost, unpublished R-Matrix (1995). New data includes collision strengths from Tayal *ApJS* **195**, 12 (2011) & A. Giunta, unpublished DW (2014).

N II term resolved line ratios

- 404/399 line ratio can be used to constrain the **temperature** (for a given density)
- 404/402 line ratio provides a measure of the **density** (for $n_e > 5 \times 10^{13} \text{ cm}^{-3}$)
- Self consistent pair of measured ratios is expected for localised N II emission in PFR

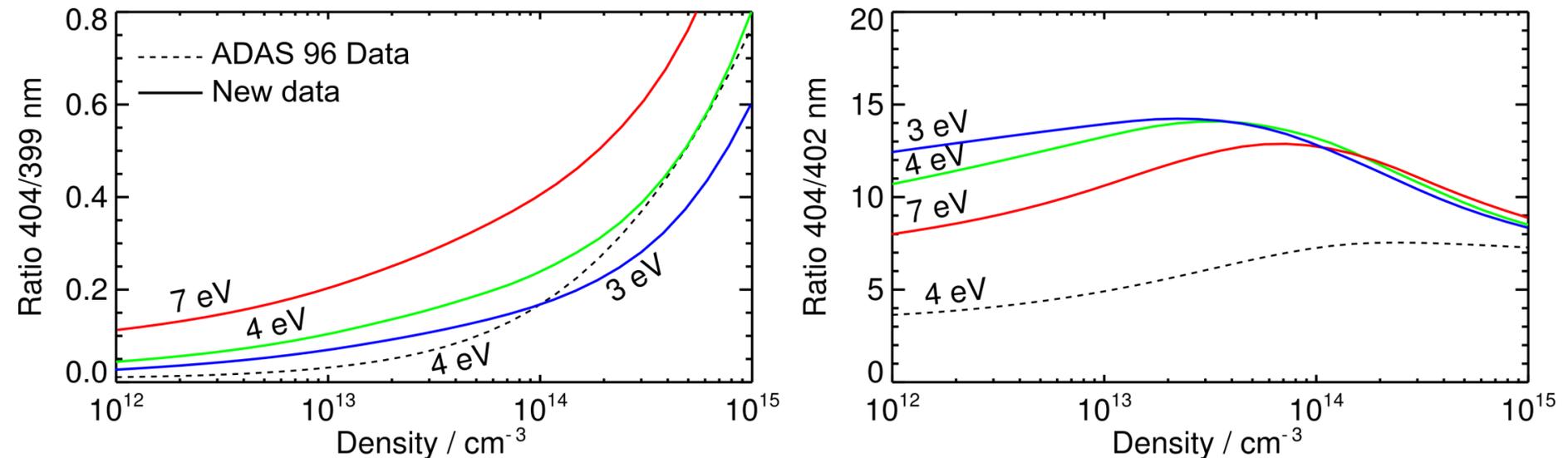


Figure 7: ADAS 96 data calculated using collision strengths from Stafford et al. *Mon. Not. R. Astr. Soc.* **268**, 816 (1994) & R. Frost, unpublished R-Matrix (1995). New data includes collision strengths from Tayal *ApJS* **195**, 12 (2011) & A. Giunta, unpublished DW (2014).

ADAS Feature Generator (AFG)

- ADAS code `split_multiplet` returns the relative intensity of lines in an LS multiplet
 - reasonable for line identification, but not always accurate for feature generation
- Level resolved intermediate coupling (IC) calculations available in ADAS
 - no difference between N II IC and LS intensities, but accurate for feature generation

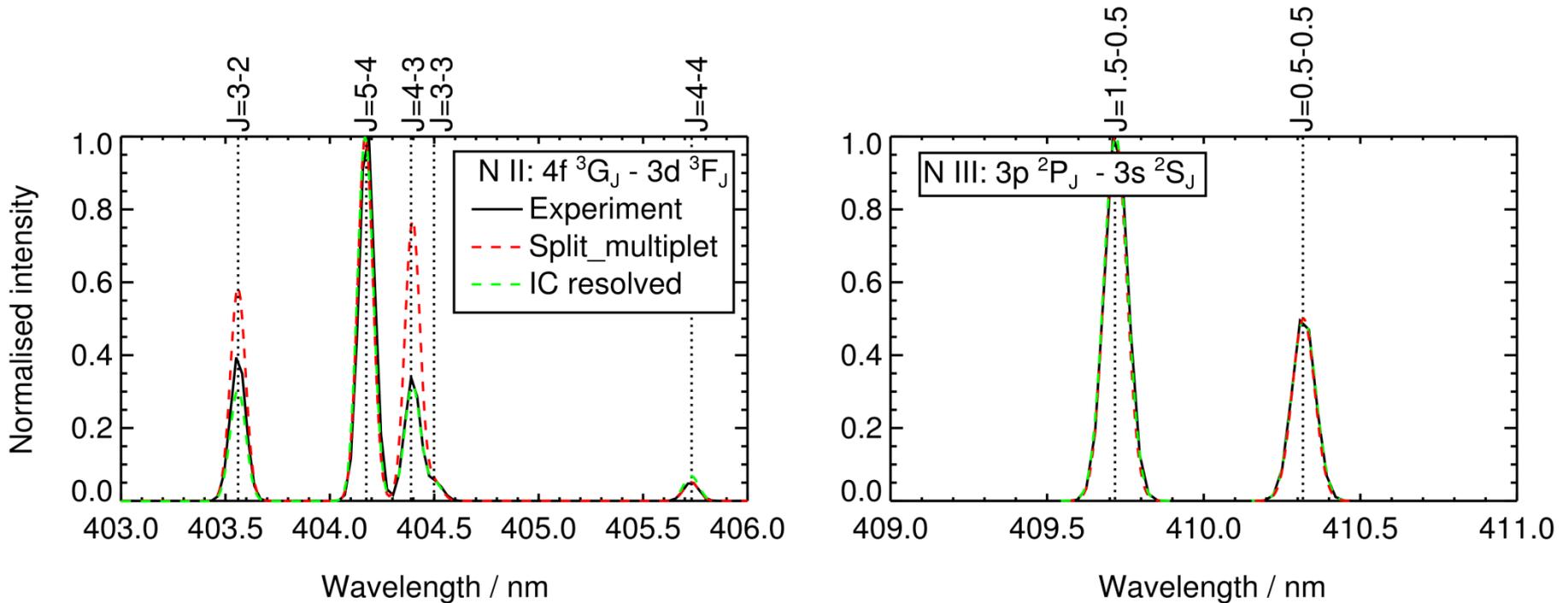


Figure 8: Experimental values represent the fitted Gaussian lines from AUG #33285. `Split_multiplet` is based on an algorithm by Condon and Shortley (Chap 9, Sec. 2, Eqn. 2a and 2b).

Framework for Feature Synthesis (FFS)

- Model of spectra includes N II AFG coupled with Gaussian & Voigtian line shapes, and the background continuum
- Least squares approach using the FFS provides fitted T_e , n_e and line integrated N^{1+} concentrations

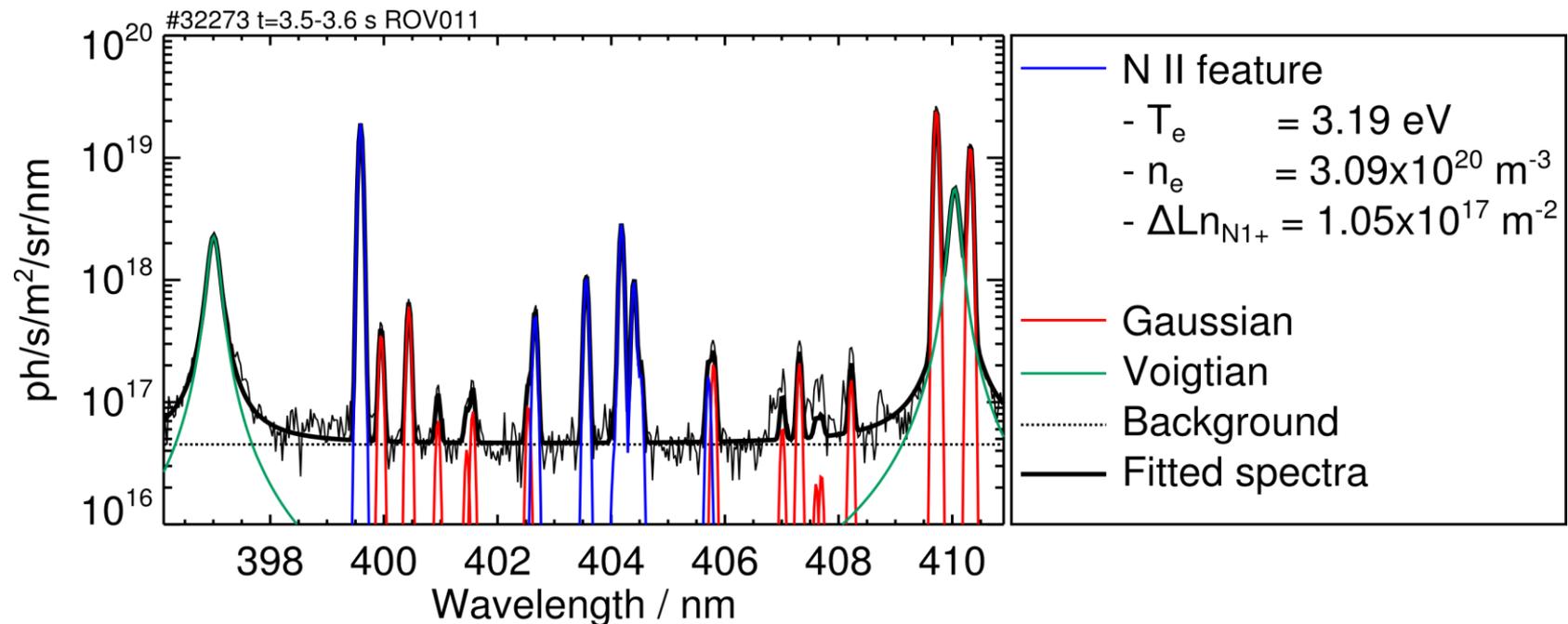


Figure 9: Model file for FFS fit includes AFG feature for N II, a multi-Gaussian parameterisation for other impurity lines, and a Voigtian parameterisation for both Balmer lines.

Framework for Feature Synthesis (FFS)

Good agreement between fitted and simulated T_e and n_e

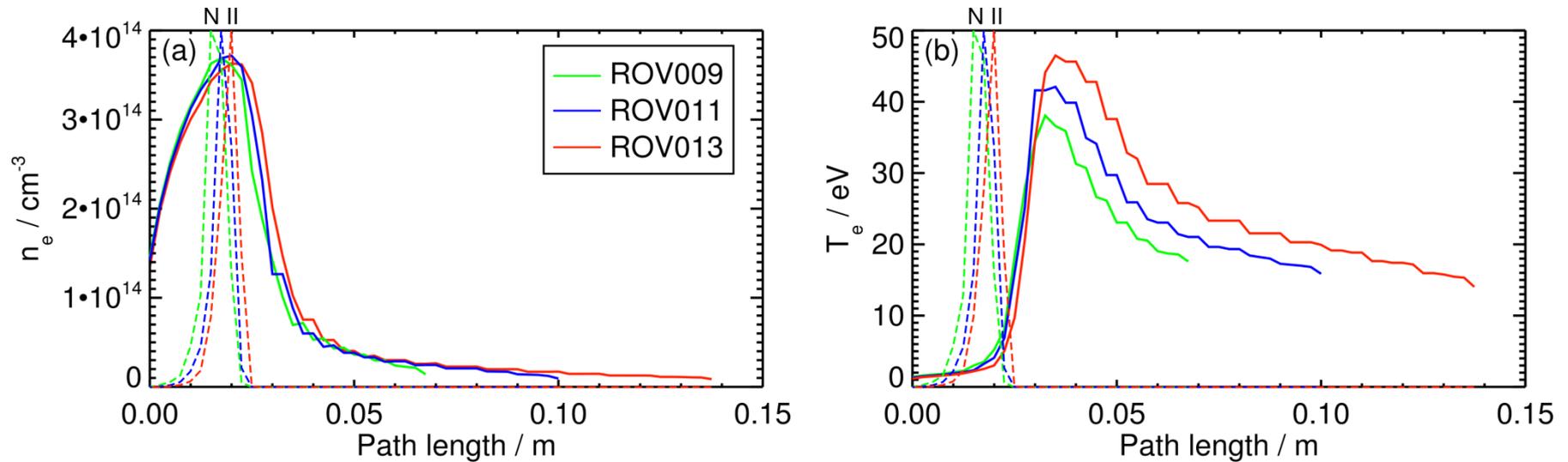


Figure 5: Sightline profiles of (a) n_e and (b) T_e as a function of the path length. Normalised N II emission is shown by the dashed lines to indicate the emitting regions.

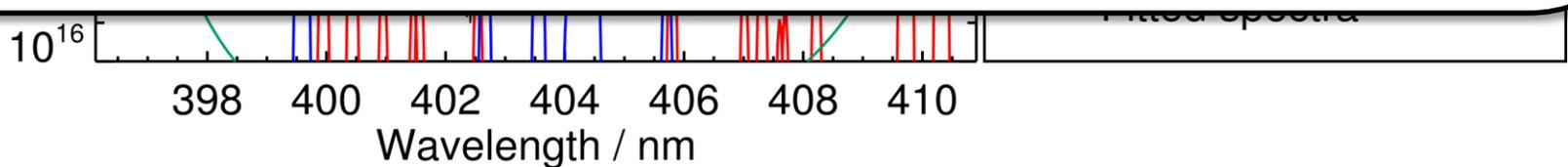


Figure 9: Model file for FFS fit includes AFG feature for N II, a multi-Gaussian parameterisation for other impurity lines, and a Voigtian parameterisation for both Balmer lines.

Temporal behaviour of fitted parameters

AUG discharge #32932 ROV011 $t_{res}=0.01$ s

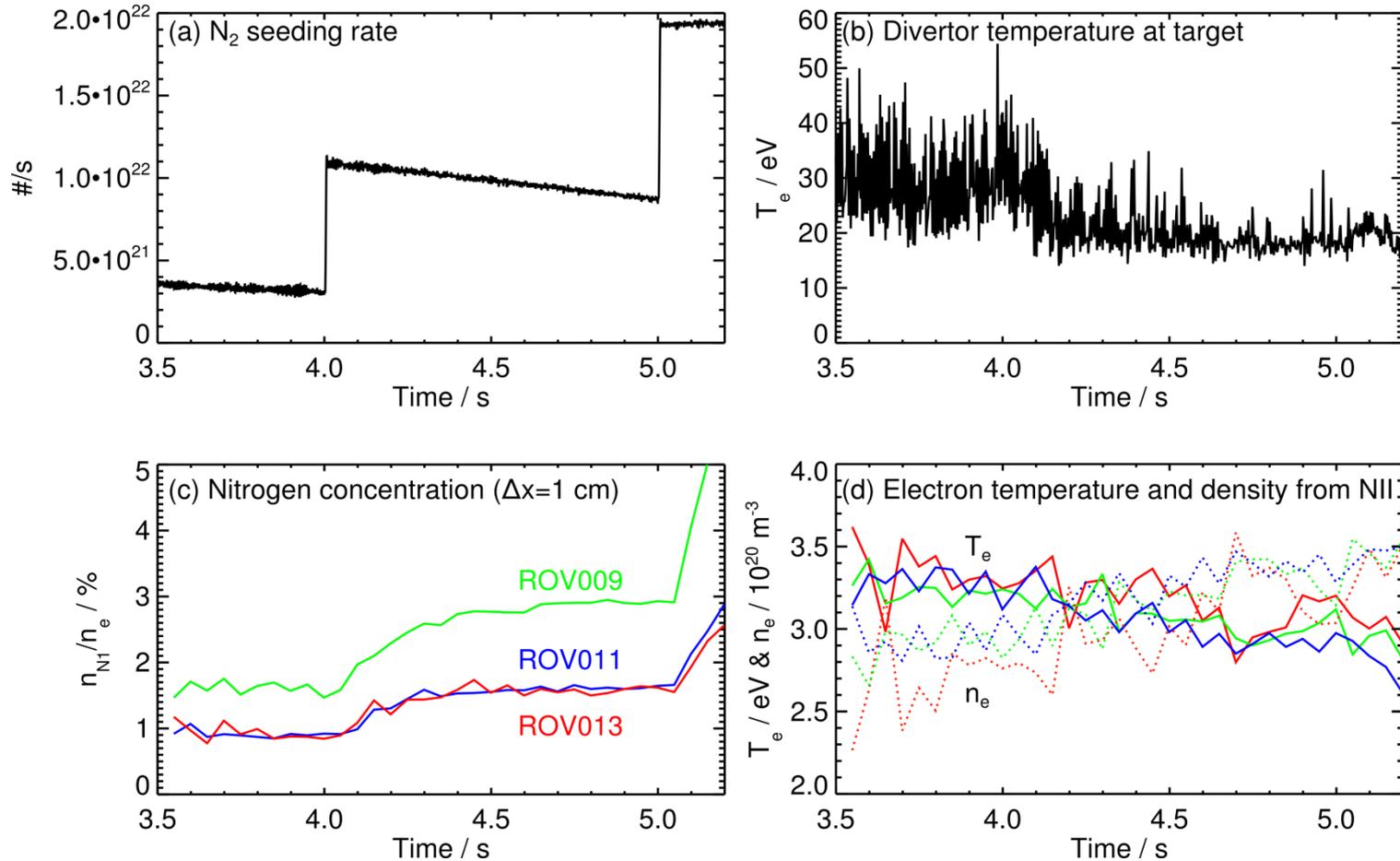


Figure 10: Fitted parameters for #32932. Time resolution has been increased to average over ELM periods. Further analysis should only model inter-ELM period (see appendix).

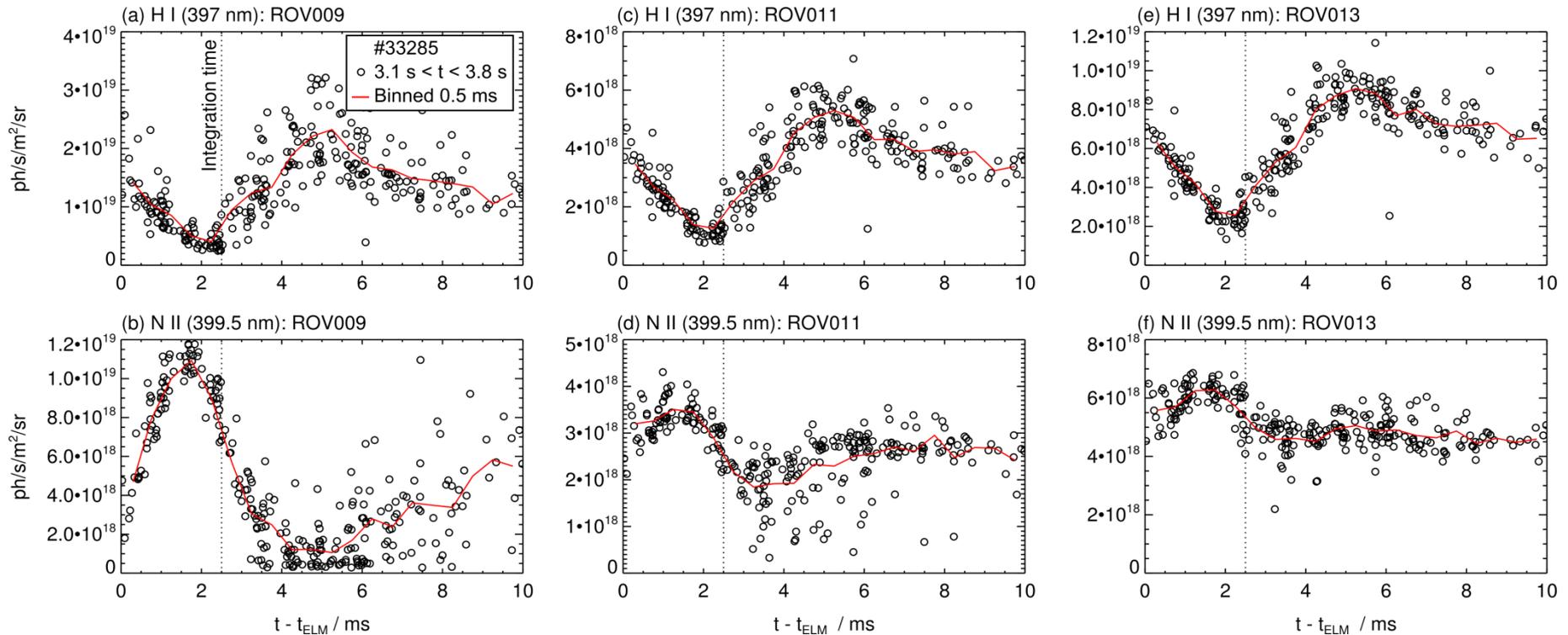
Summary

- 1) Narrow spectral range used for H_{δ} and H_{ϵ} Stark broadening measurements also provides a number of impurity lines useful for analysis
- 2) SOLPS simulations for AUG suggest that N II (and N III & N IV) emission is localised to PFR during attached divertor conditions
- 3) Atomic modelling of two line ratios (from three N II lines) provides a measure of the electron density and temperature, and therefore also the line-integrated N^{1+} concentration
- 4) Absolute values of fitted parameters agree well with SOLPS predictions in the PFR, while the temporal behaviour of fitted parameters respond accordingly to changes in N_2 seeding rate and temperature at the divertor target

Thanks for listening, any questions?

Appendix

Distinguishing between ELM and inter-ELM



- Region of hot plasma during the ELM period ($t - t_{\text{ELM}} = 0 - 2$ ms)
- Plasma cools between $t - t_{\text{ELM}} = 2 - 5$ ms
- Plasma relaxes back to initial condition $\sim 5 - 10$ ms after ELM
- **N II emission closer to target (ROV009) most sensitive to ELMs**